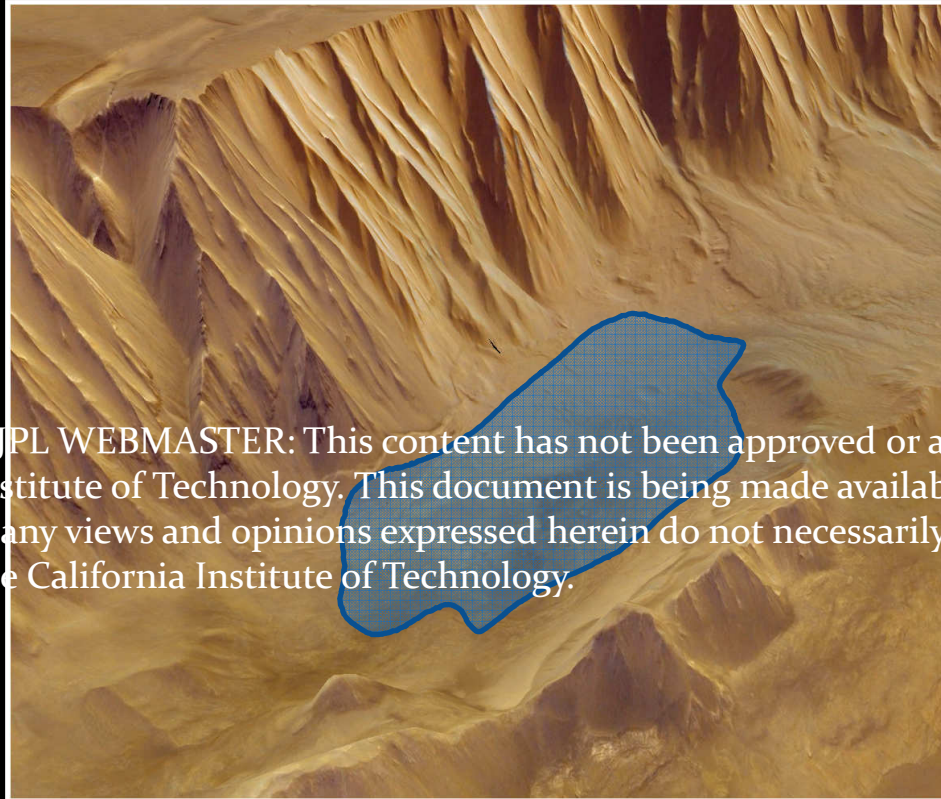


In Situ investigation of the Southwestern Melas Basin



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Rebecca M. E. Williams, Catherine M. Weitz,

Peter M. Grindrod, Joel Davis



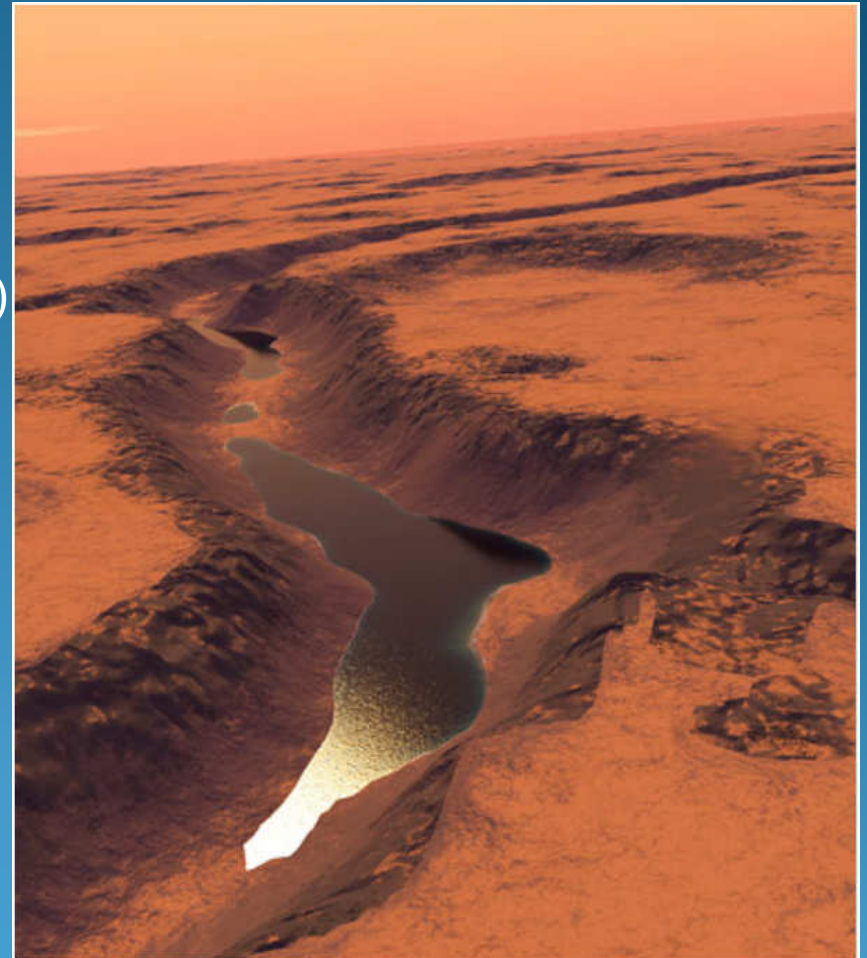
Catherine Quantin-Nataf, Gilles Dromart



Seeking Signs of Life

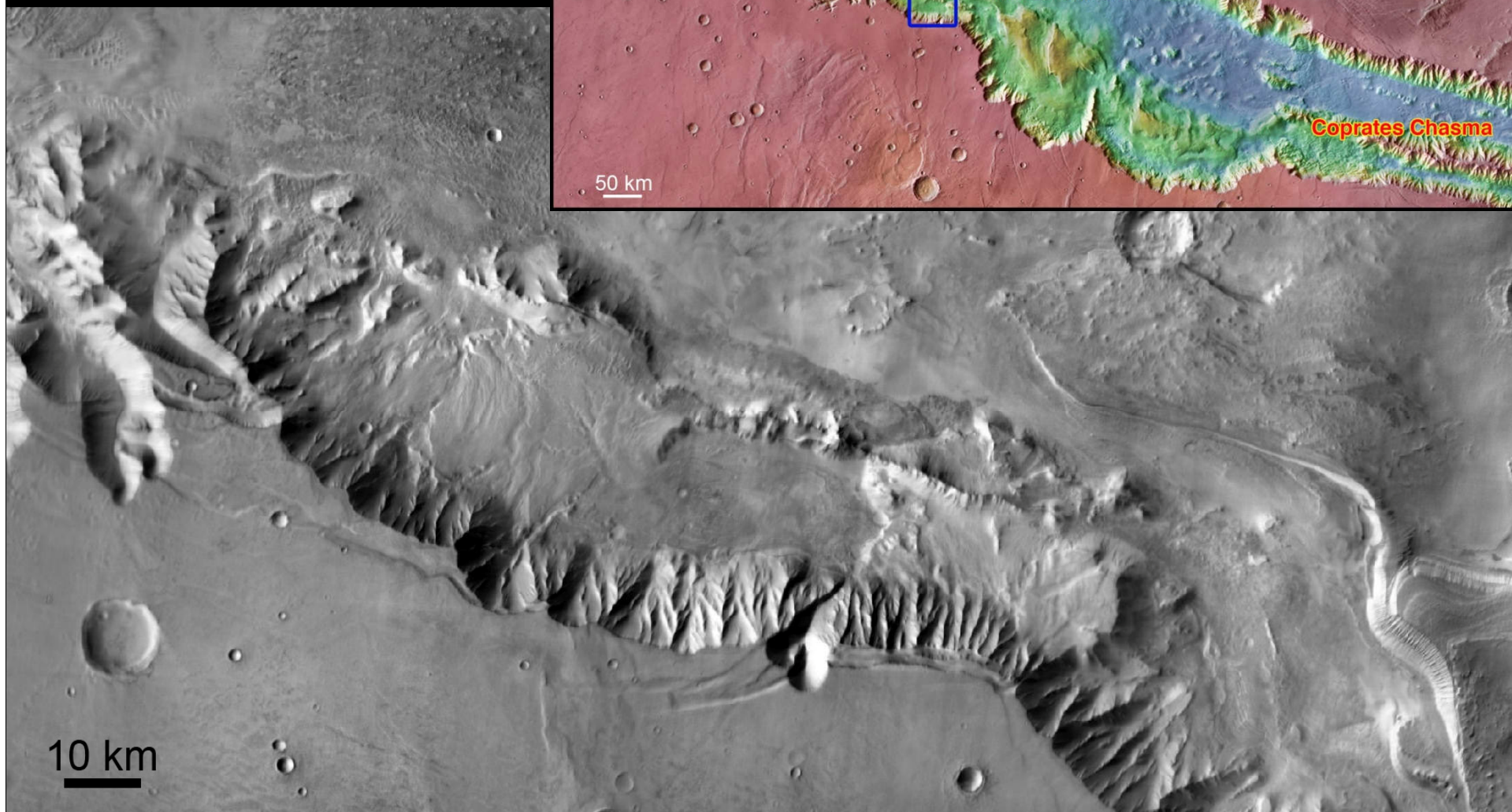
2020 Mars Rover Science Objectives

- A. Habitable ancient environment
 - Evidence of aqueous activity
- B. Biosignature preservation potential (BPP)
 - Depositional environment (preferably lacustrine)
- C. Scientifically selected, well-documented samples
 - Geologic and stratigraphic context
 - Accessibility
- D. Pave the way for human exploration
 - Achievable at all landing sites

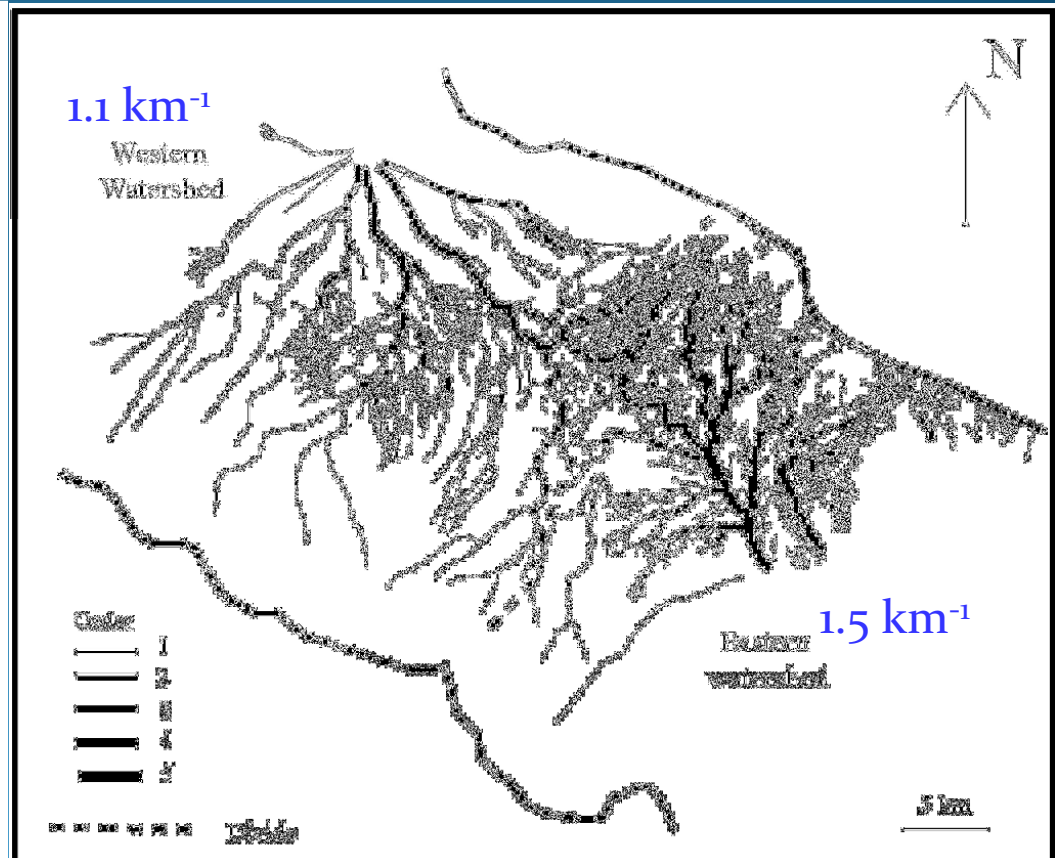
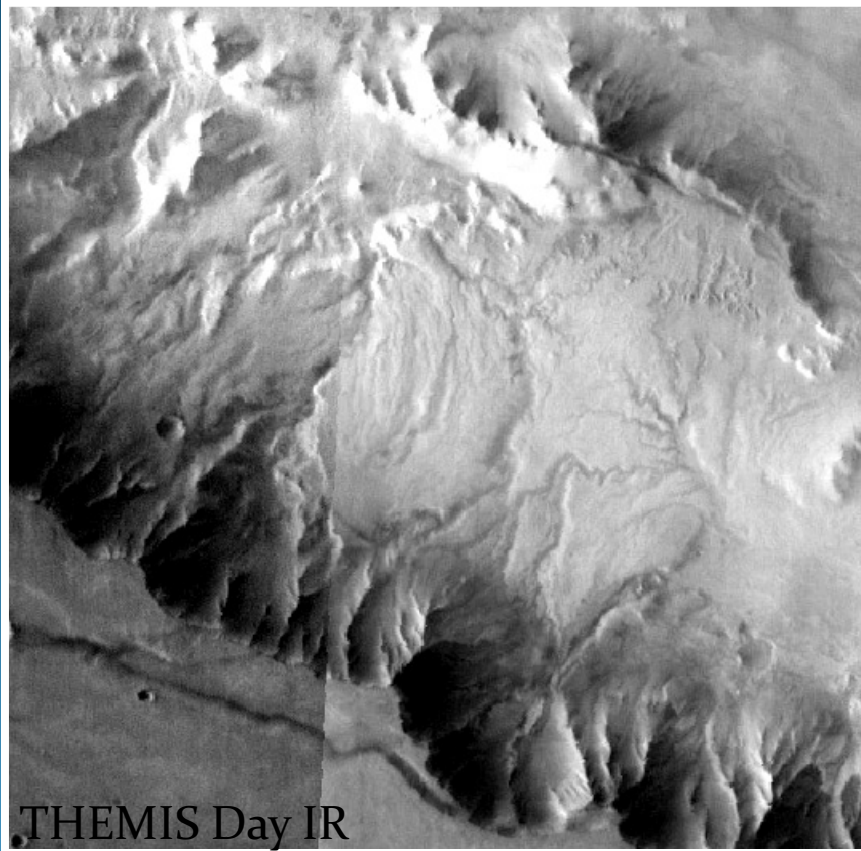


(National Geographic)

Regional Setting



Western Drainage Basin



(Quantin et al., *JGR*, 2005)

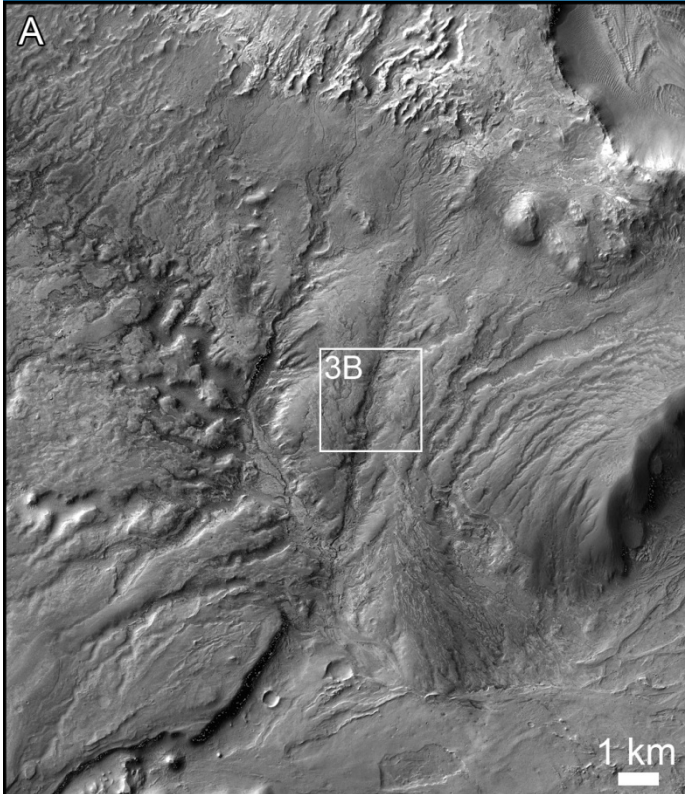
Maximum VN Drainage Density 0.14 km⁻¹ (Hynek et al., 2010)

VN Maturity implies formation timescales of thousands of years.

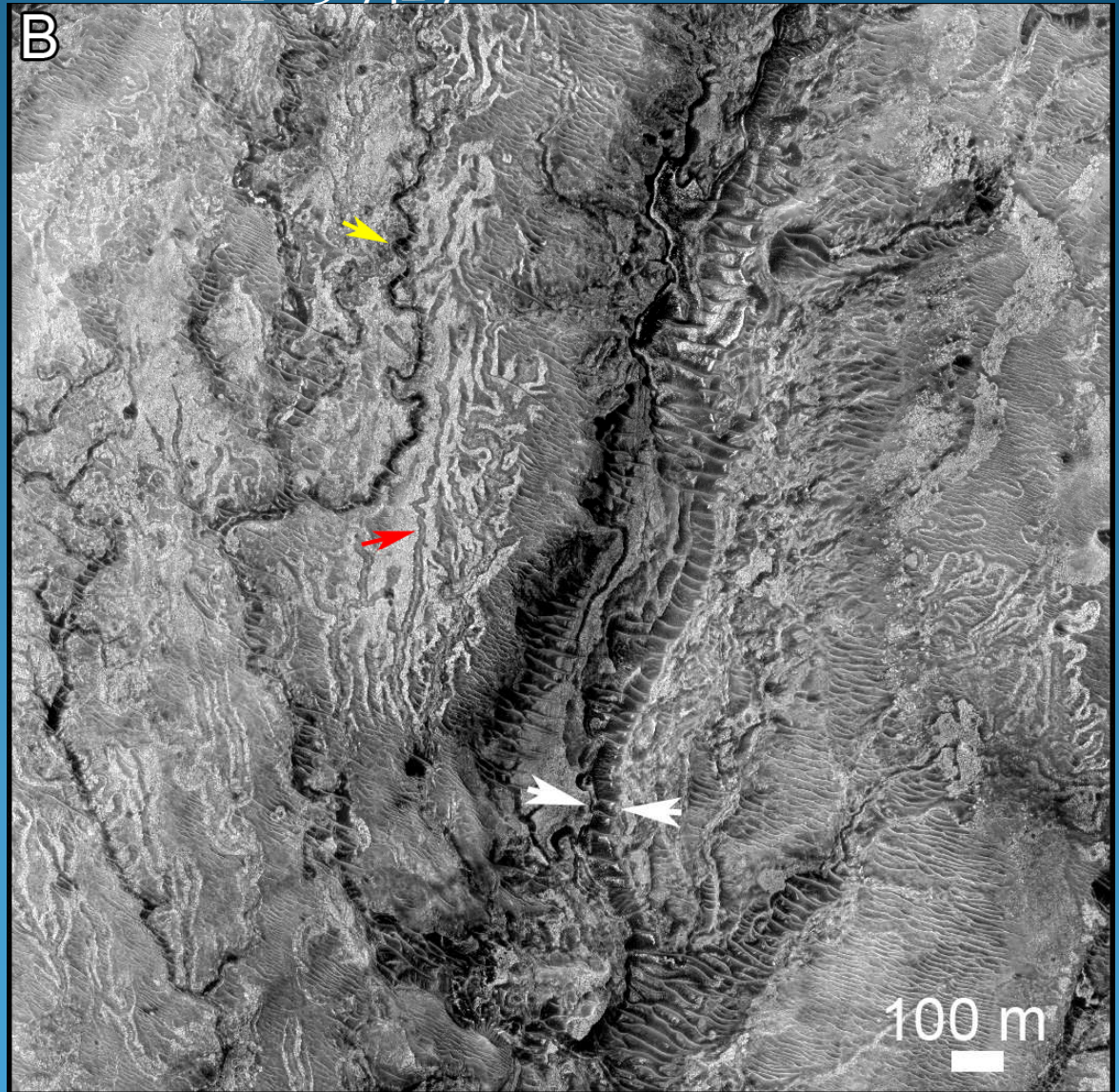
Western Drainage Basin

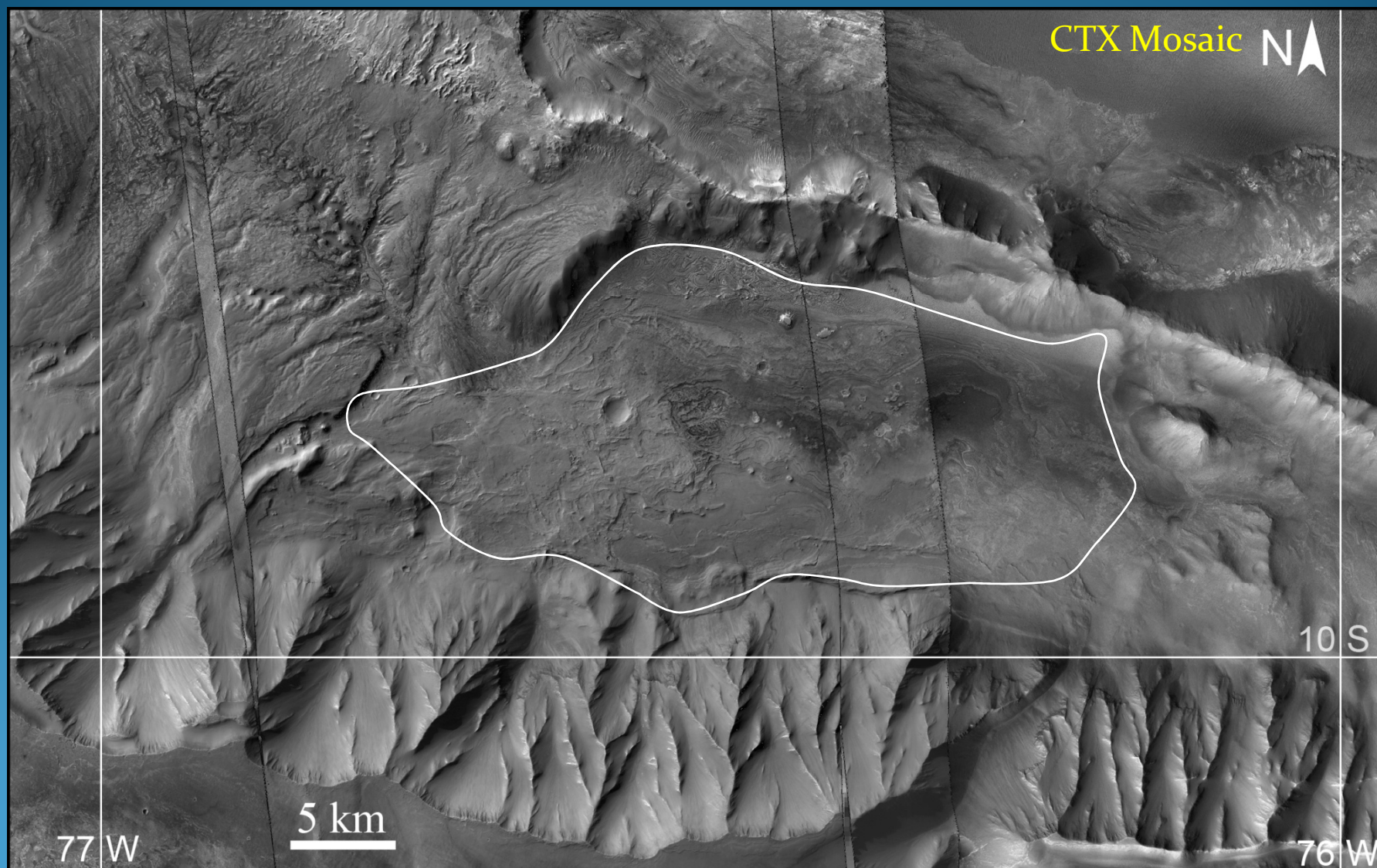
HiRISE PSP_005874_1700

CTX P20_009025_1704



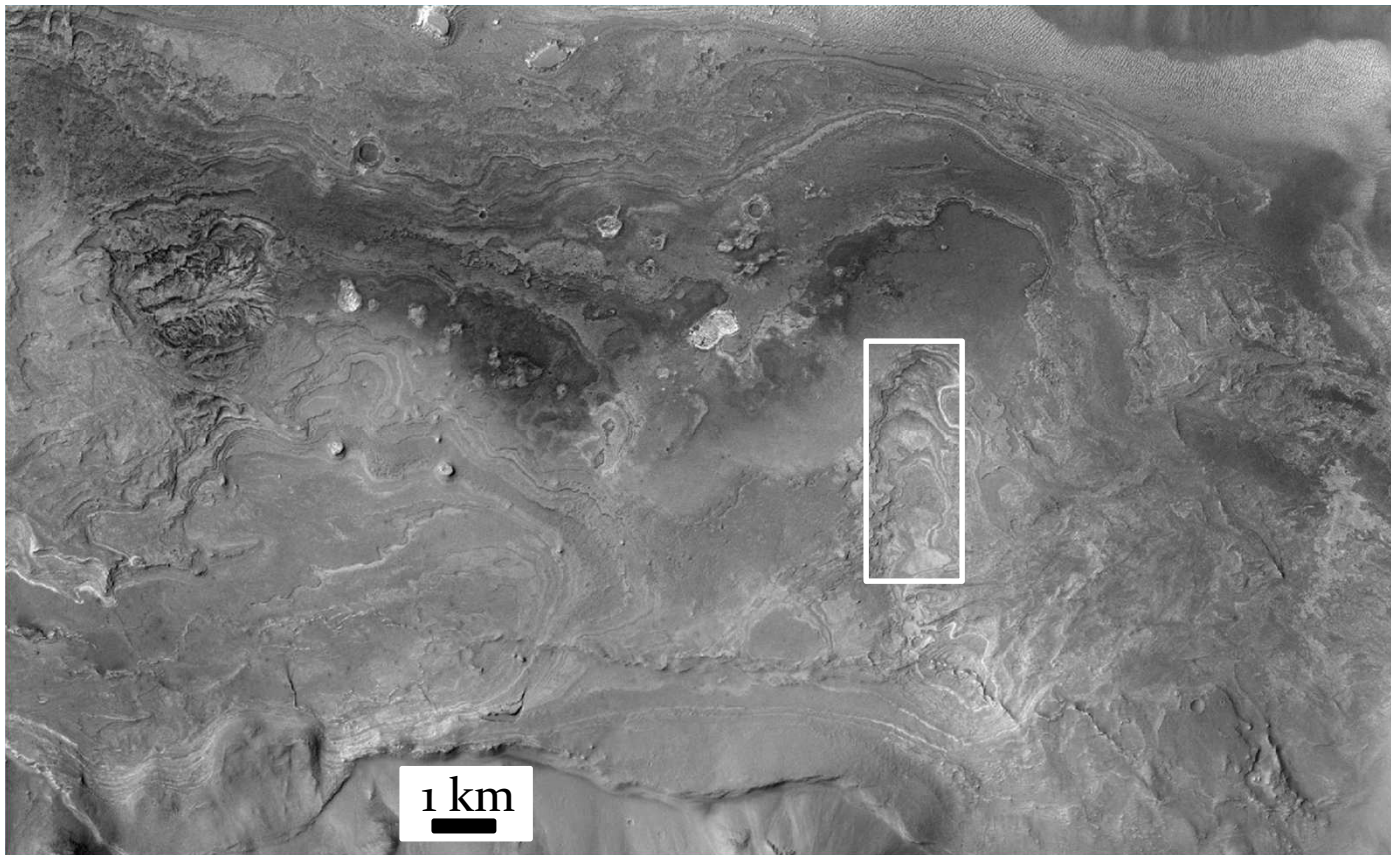
Early parallel drainage
Later: Sinuous incision

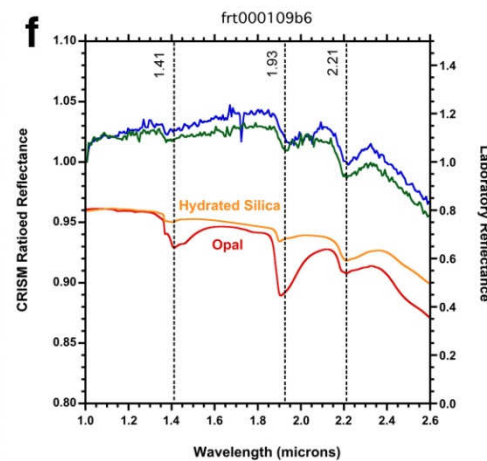
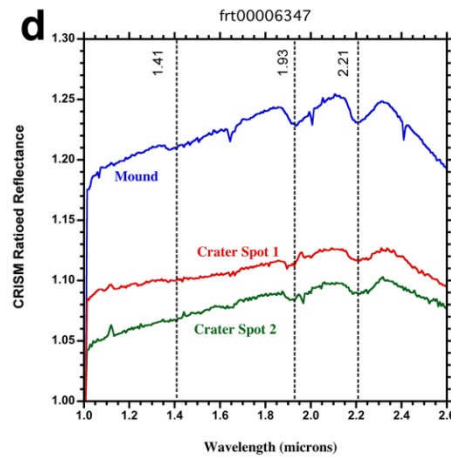
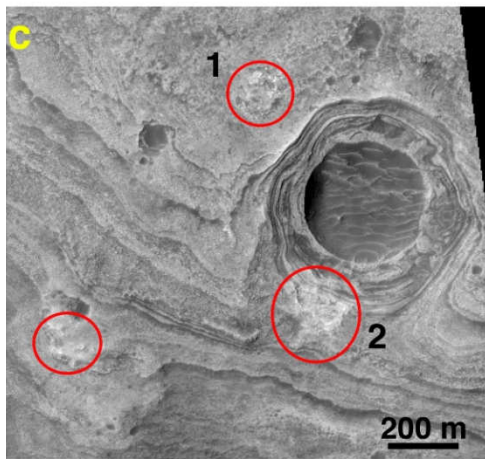
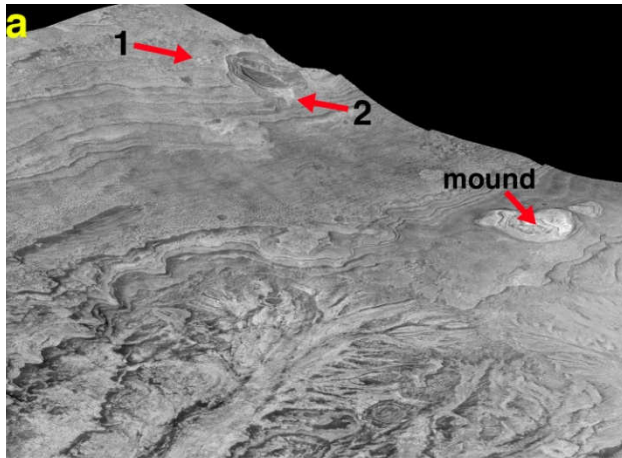




Layered Beds

Hypotheses: lacustrine, aeolian, volcanic

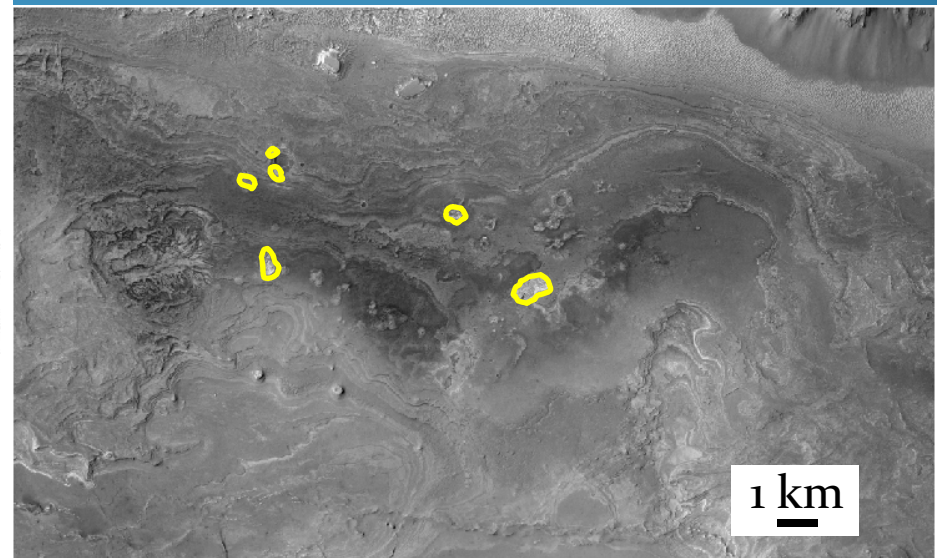




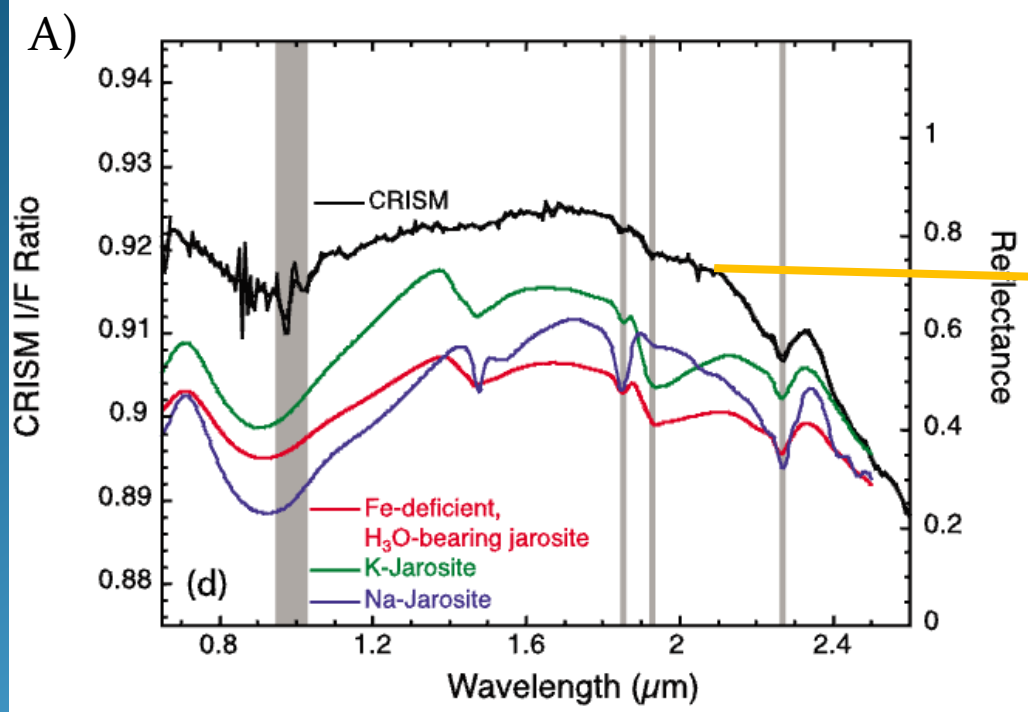
Hydrated Silica/Opal

- Associated with fresher, brighter exposures of pre-existing beds
- Not tied to one bed or elevation
- Alteration of sediments by unknown processes

(Weitz et al., 2014)

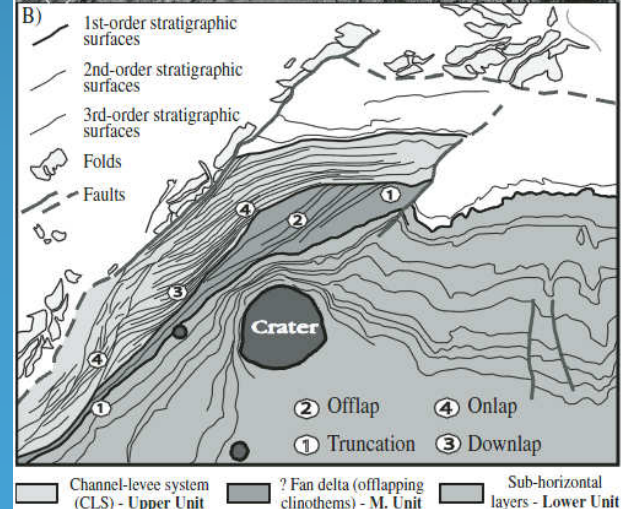
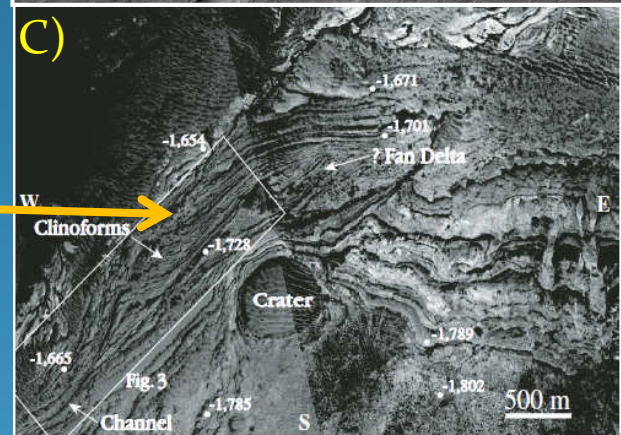
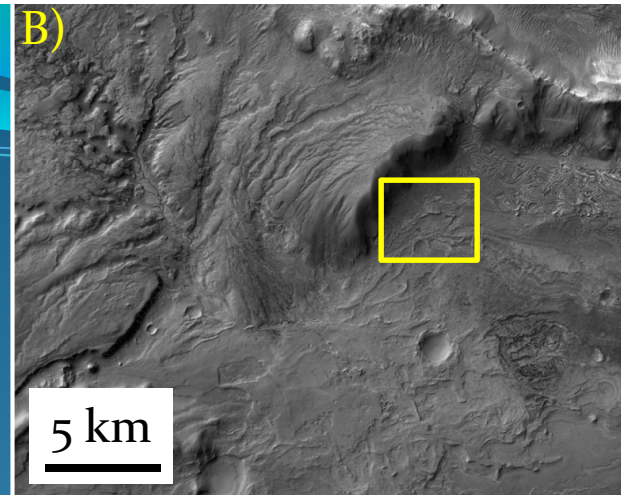


Sulfates—Jarosite

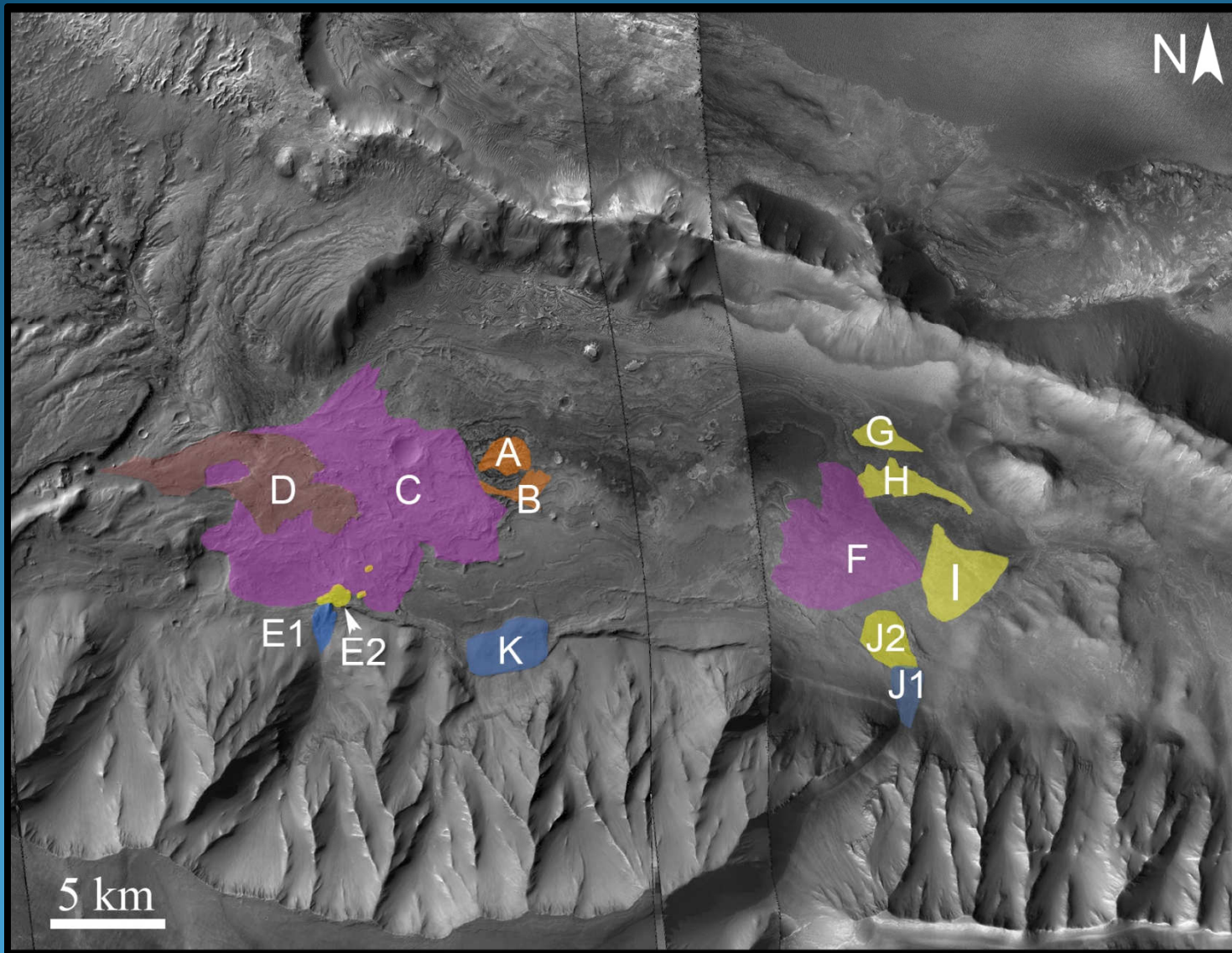


(Metz et al., 2009)

(Dromart et al., 2007)



Depositional Environments



Subaerial

Brown = Debris flows

Blue = Landslides

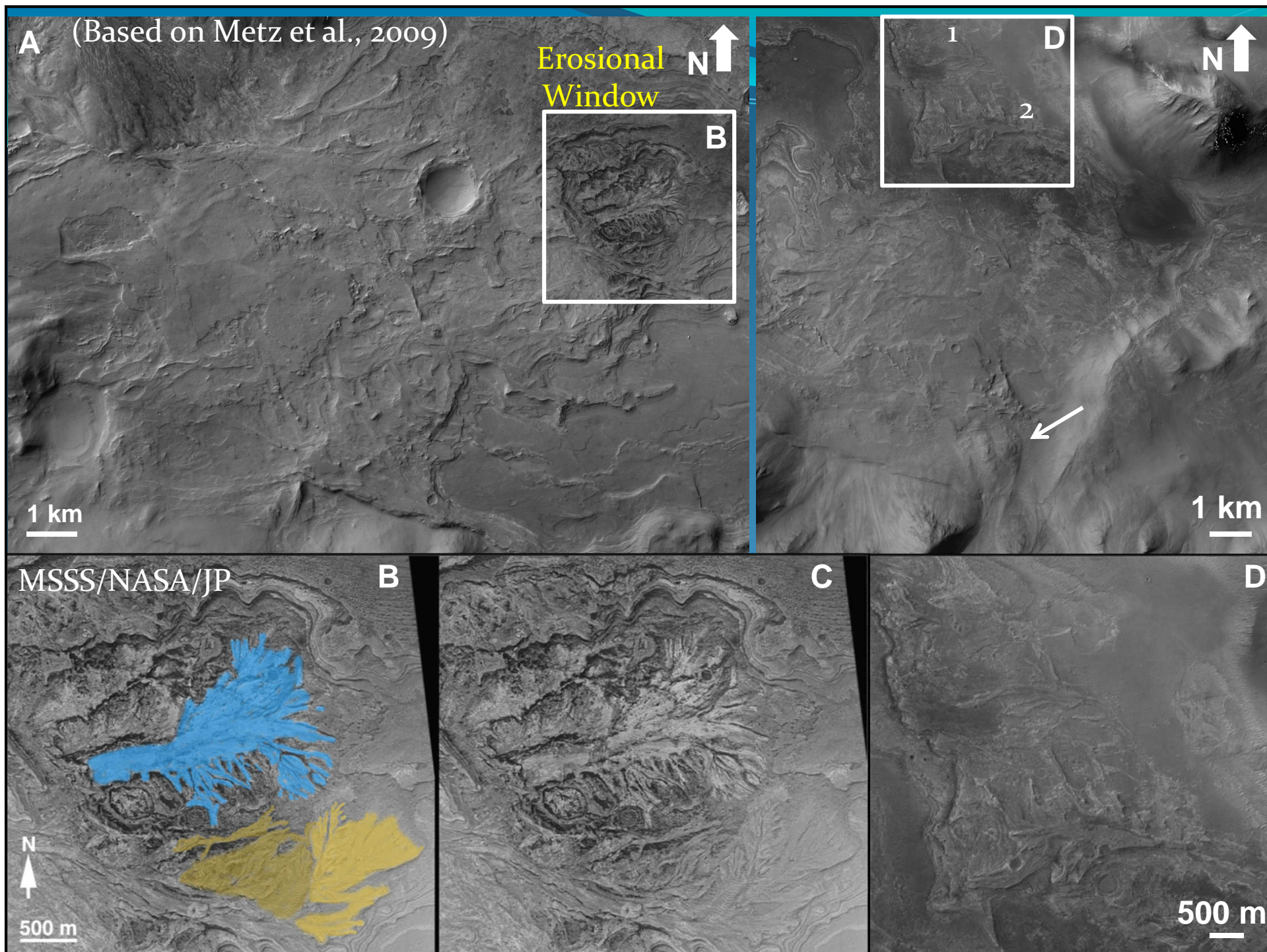
Subaqueous

Purple = Fan/deltas

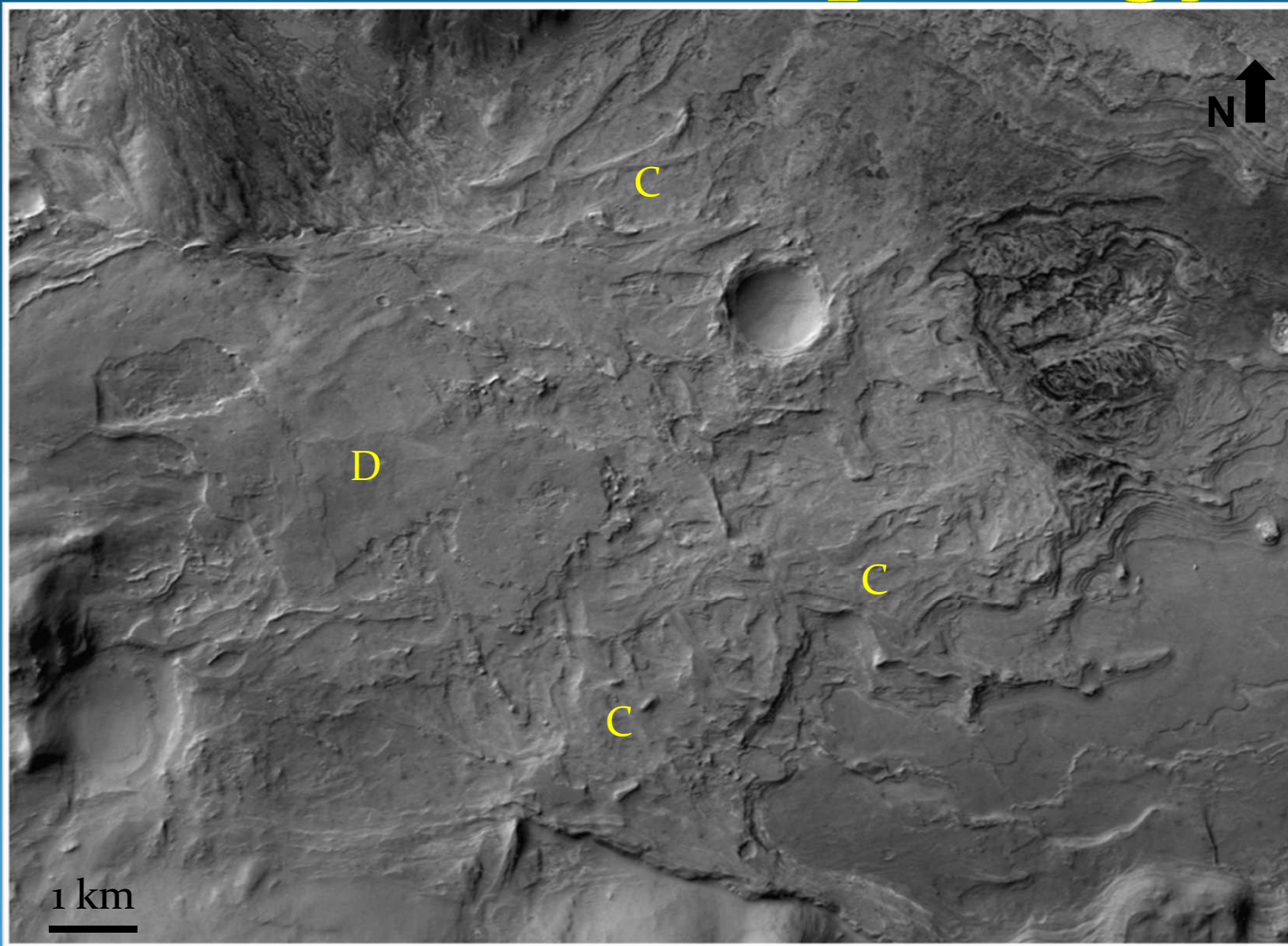
Yellow = Deltas

Orange = Deep
sublacustrine fans

(Williams and Weitz,
in review)

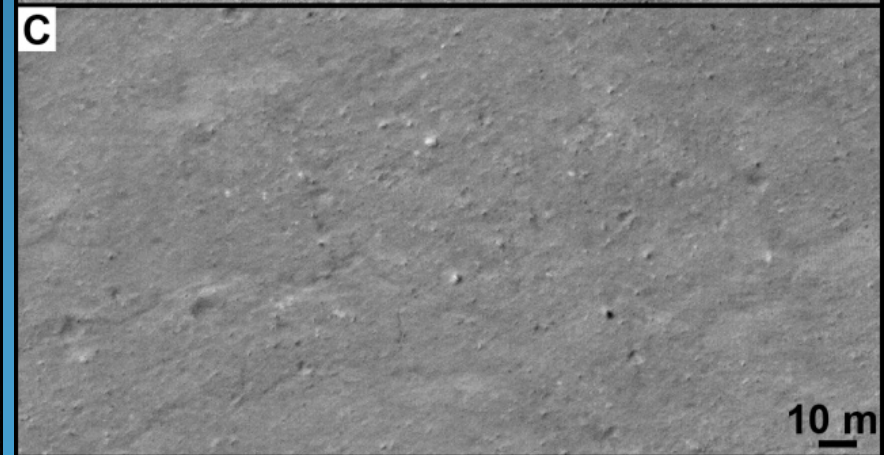
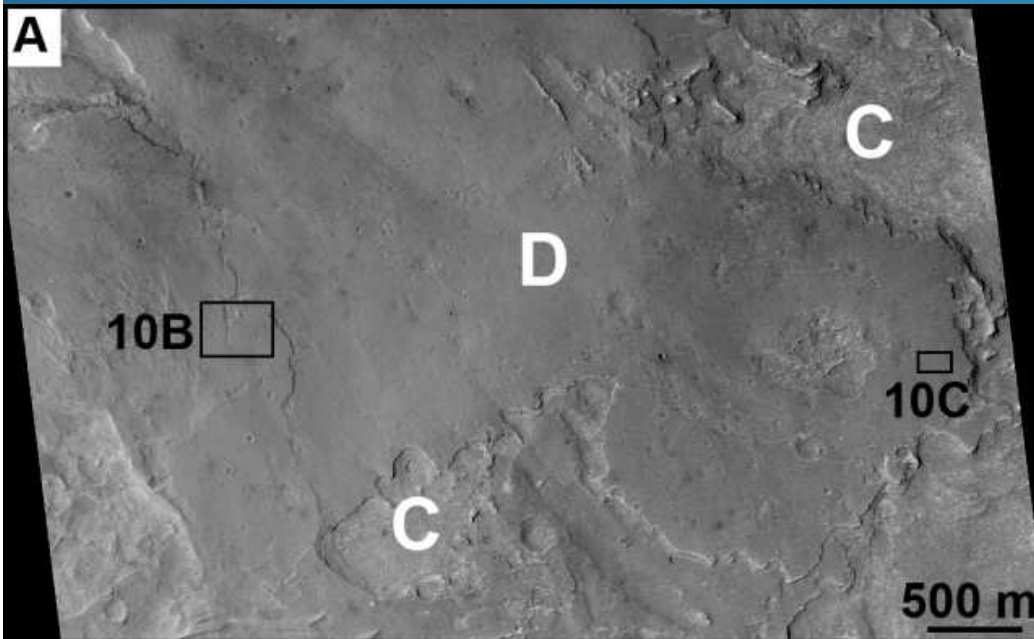
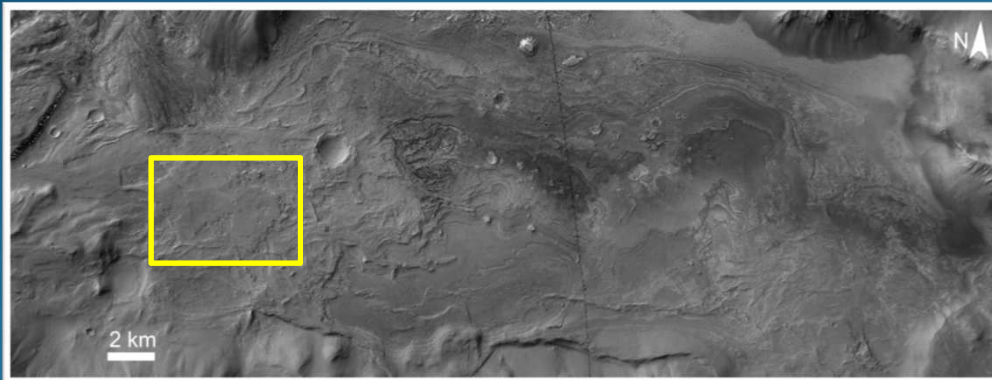


Fans C & D: Morphology



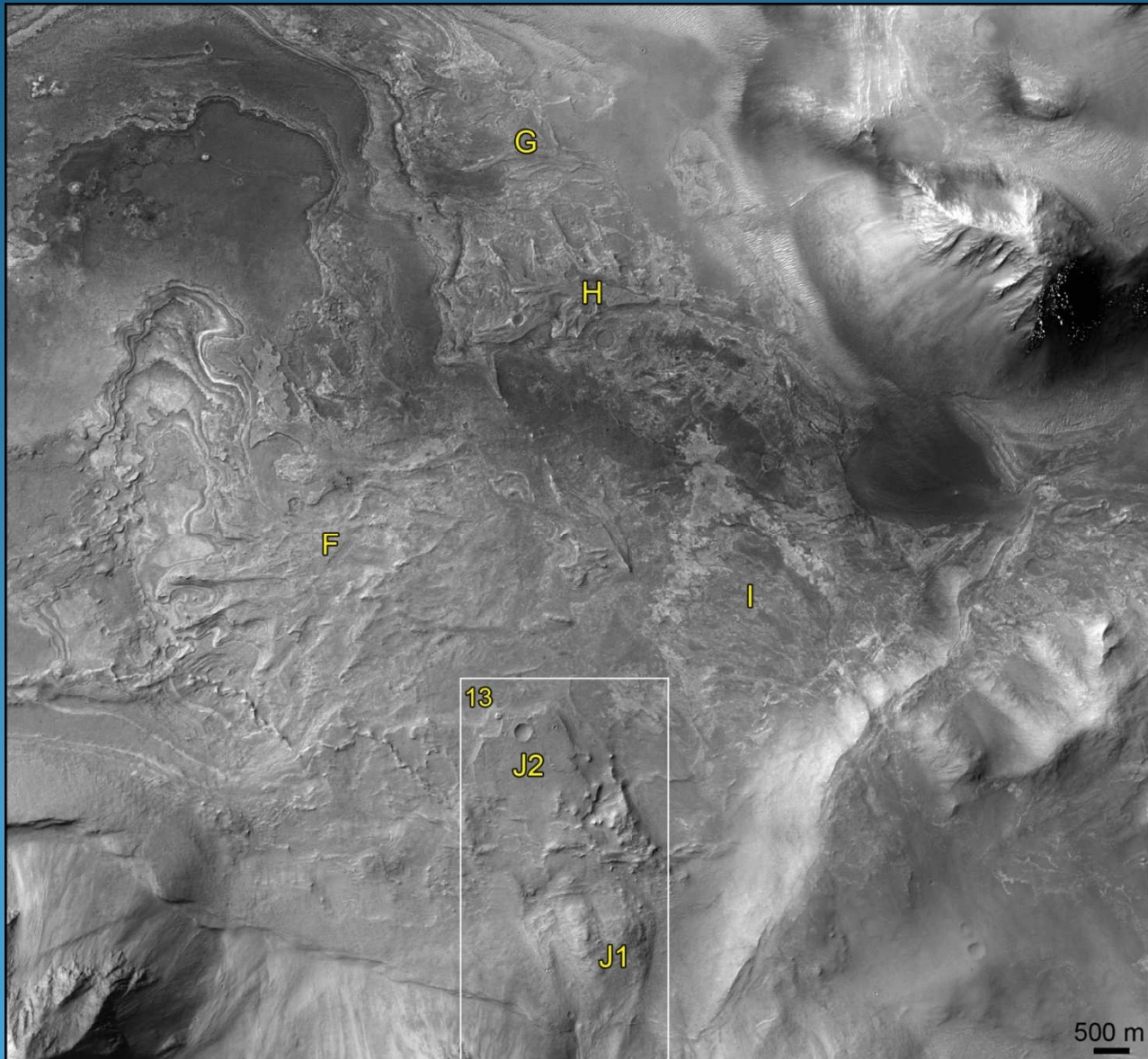
(Williams and Weitz, *in review*)

Fans D: Morphology

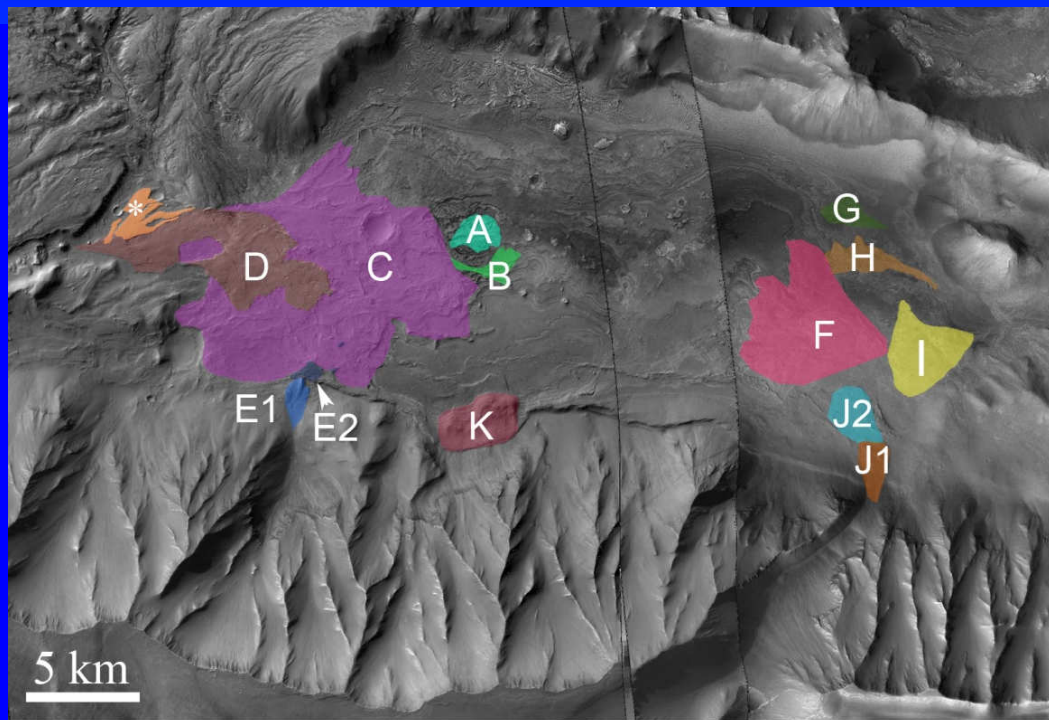
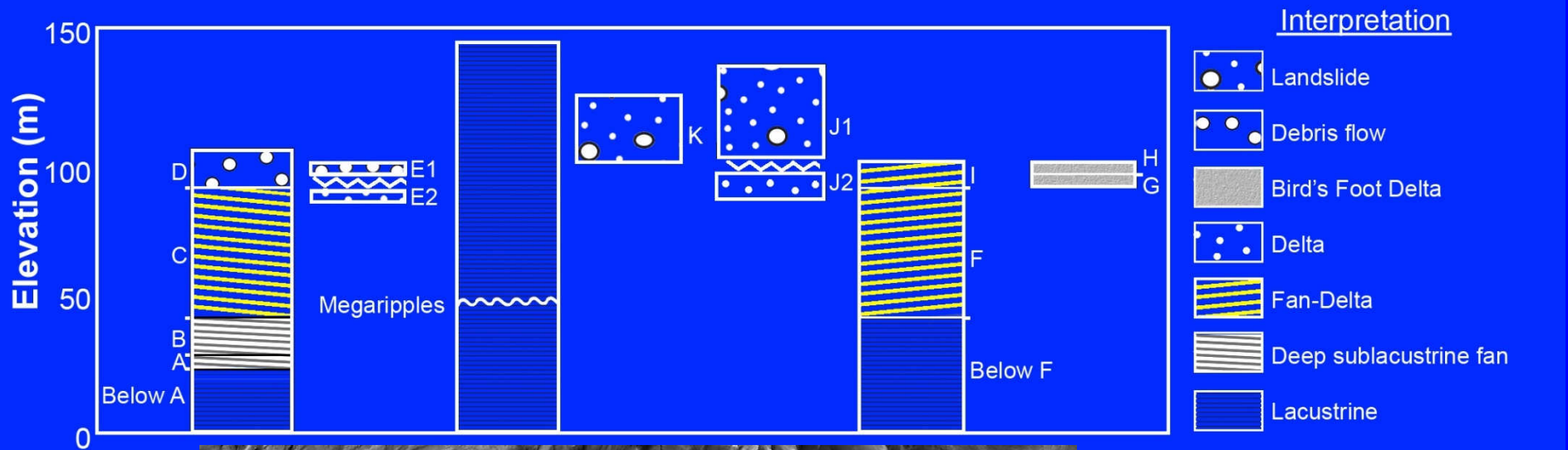


(Williams and Weitz, *in review*)

Eastern Fan Complex



Stratigraphic Relationships



(Williams and Weitz, *in review*)

Timescales

1) Use average discharge ($\sim 30 \text{ m}^3/\text{s}$), representative of persistent flow conditions

→ 100-250 years basin fill time

Only western inflow

2) Use large delta deposition rates ($\sim 0.001 \text{ km}^3/\text{yr}$)

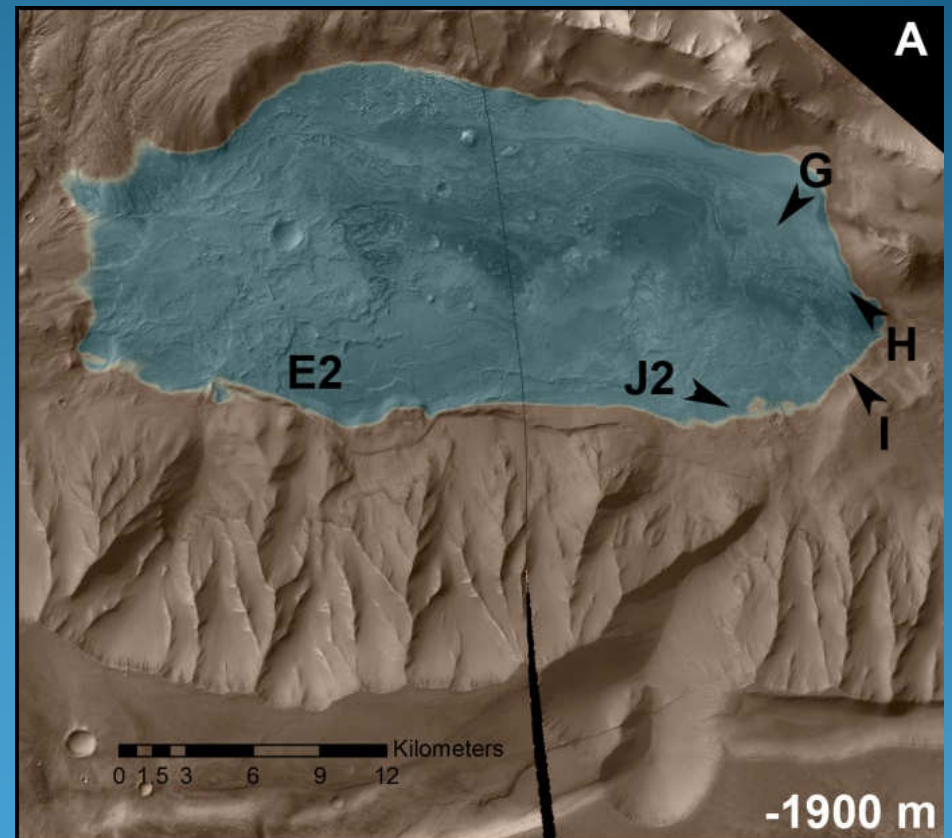
Fan C Volume 0.1 km^3

→ ~100 years Fan C formation time

3) Use lacustrine deposition rates ($0.02 - 1 \text{ cm/year}$)

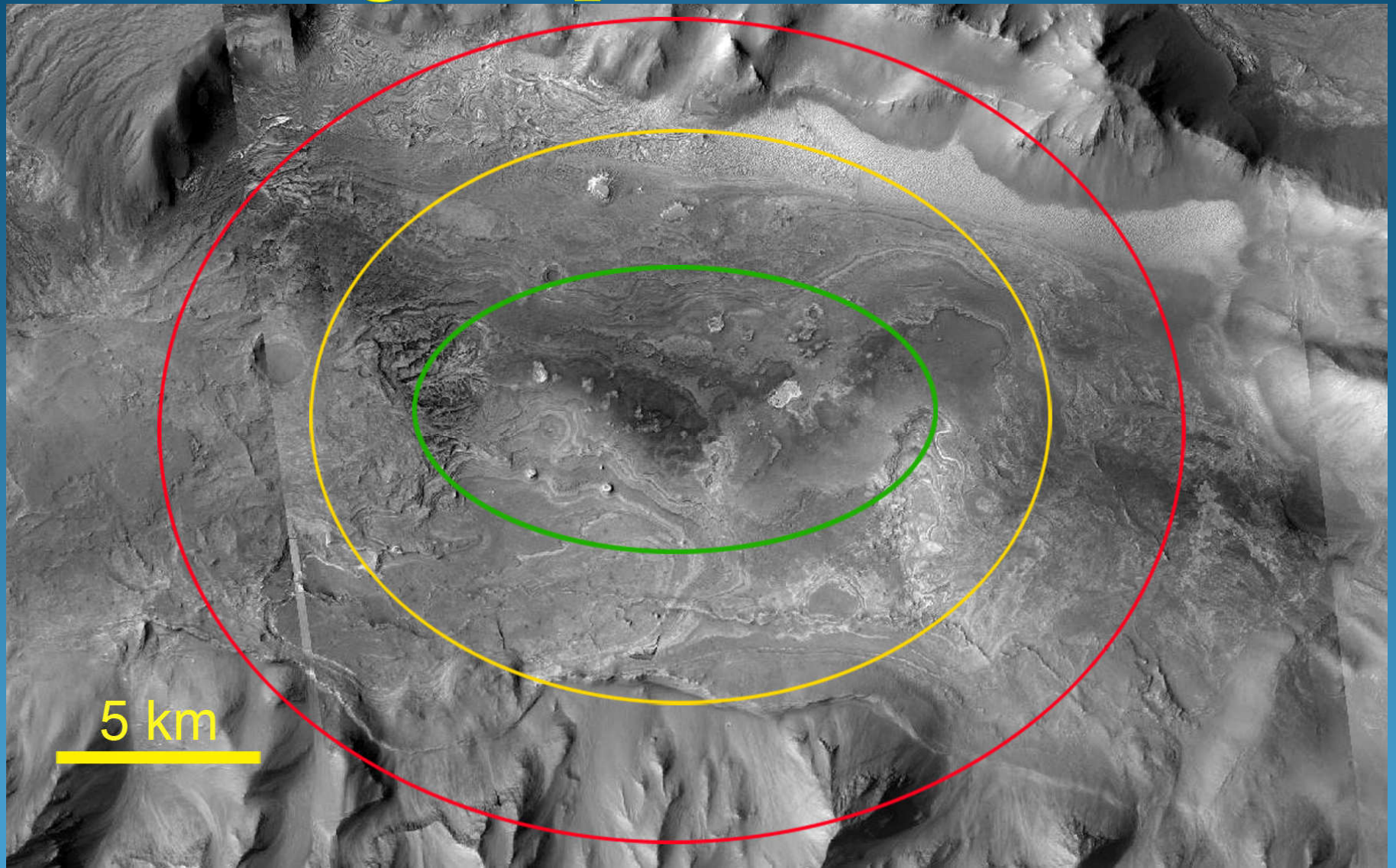
→ for 25 m layered beds below fan A

$10^3 - 10^5$ years



(Williams and Weitz, *in review*)

Landing Ellipses

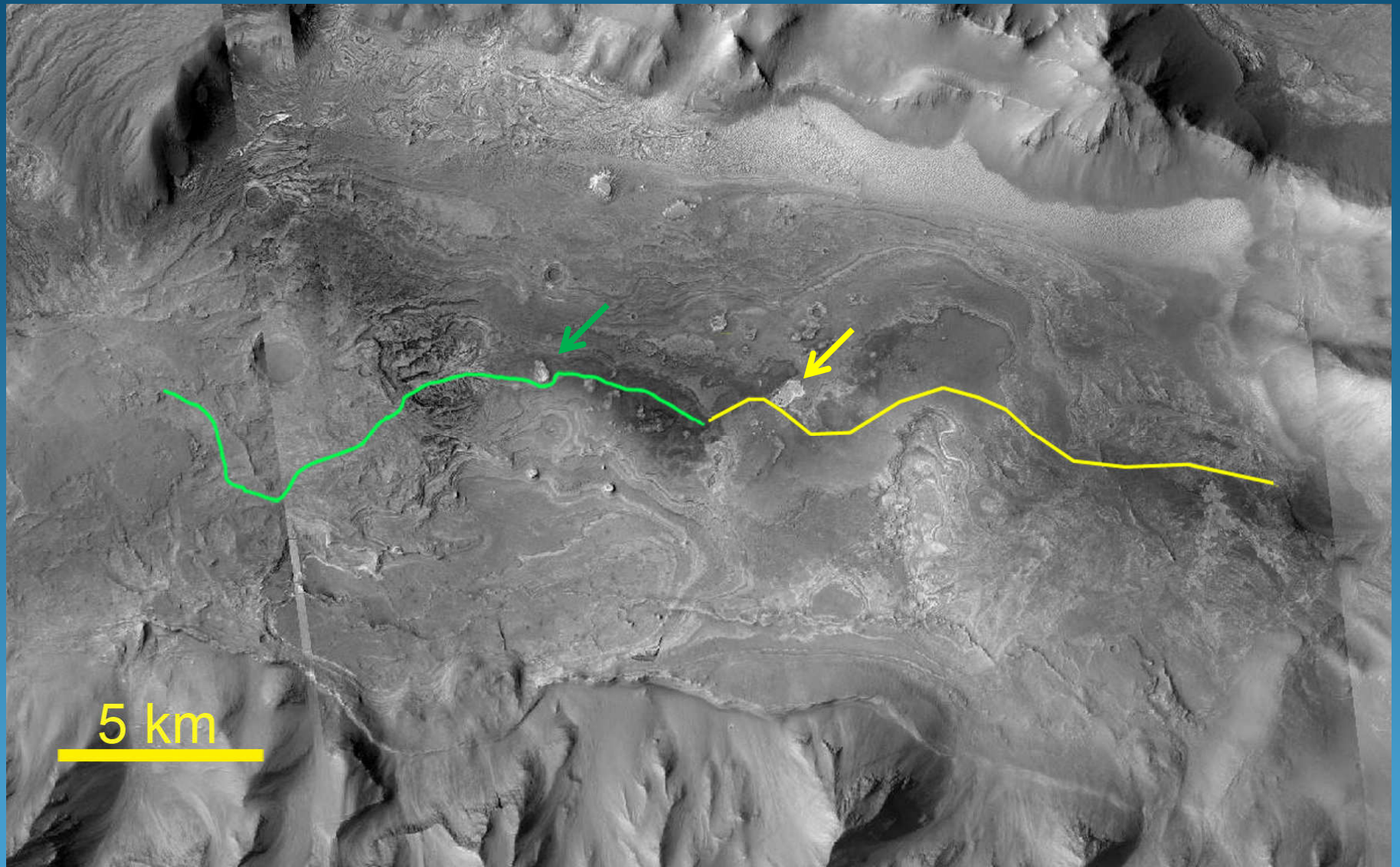


Red: 25 x 20 km

Yellow: 18 x 14 km

Green: 13 x 7 km

Potential Routes



Red: 25 x 20 km

Yellow: 18 x 14 km

Green: 13 x 7 km

Engineering Constraints

Parameter	Criteria	Compliance
Elevation	Below 0.5 km	✓ -1.850 +/- 0.100 km
Latitude	Within $\pm 30^\circ$ of the equator	✓ 9.84°S , 283.63°E
Slope	$< 20^\circ$ over 2-10 km	✓
	$< 25^\circ$ over 2-5 m	✓ $> 98\%$ for smallest ellipse
Relief	< 100 m over 1-1,000 m	Few locations within ellipse exceed relief constraint.
Rock Abundance	Probability of rock 0.55 m tall in area of 4 m^2 is $< 0.5\%$	HiRISE verification needed
Thermal inertia	$> 100 \text{ J m}^{-2} \text{ s}^{-0.5} \text{ K}^{-1}$	✓
Albedo	< 0.25	✓

See supplemental slides for more details

Scientific Rationale for SW Melas

- Well defined geologic and stratigraphic context
- Long-lived lacustrine history exposed
- Scientific targets are distributed across entire ellipse
- Exhumed terrain means young exposure age
 - protection from irradiation
- Diversity of sample targets within landing ellipse
 - Lacustrine deposits
 - Deltaic deposits
 - Alluvial fans
 - Debris flows
 - Hydrated silica
 - Sulfates
 - Potential Rock Targets
 - Volcanic ash deposit
 - Transported materials from uplands
 - Impact breccia
 - Igneous rocks

Challenges for SW Melas

- Requires range trigger to shrink ellipse size; would benefit from TRN
- Minor violation of relief constraints
- No *in situ* lava flow
- Detailed examination of rock abundance and traversability is needed

Supplemental Slides

Threshold Geological Criteria

SW Melas

1. Presence of subaqueous sediments or hydrothermal sediments (equal 1st priority),

✓ Deltaic & lacustrine

OR

hydrothermally altered rocks or low-T fluid-altered rocks (equal 2nd priority)

✓ Hydrated silica

2. Presence of minerals indicative of aqueous phases (e.g., phyllosilicates, carbonates, sulfates, etc.) in outcrop

✓ Hydrated silica

3. Noachian/Early Hesperian age based on stratigraphic relations and/or crater counts

✓ Noachian/Early Hesperian Age

4. Access to unaltered igneous rocks as float

Uncertain

Potential Qualifying Geological Criteria

SW Melas

- | | |
|-----------------------------------------------------------------------------------------------------------------------------------|---------------------------------------------------|
| 1. Morphological criteria for standing bodies of water and/or fluvial activity | √ Deltaic & lacustrine deposits |
| 2. Assemblages of secondary minerals of any age. | √ Hydrated silica |
| 3. Presence of former water ice, glacial activity or its deposits. | Not recognized. |
| 4. Igneous rocks of Noachian age, of known stratigraphic relation, better if including exhumed megabreccia. | Inferred volcanic layers in wall rock. |
| 5. Volcanic unit of Hesperian or Amazonian age well-defined by crater counts and well-identified by morphology and/or mineralogy. | Not recongized. |
| 6. Probability of samples of opportunity (ejecta breccia, mantle xenoliths, etc.). | Ejecta breccia, 'grab bag' of catchment lithogies |
| 7. Potential for resources for future human mission | Uncertain. |

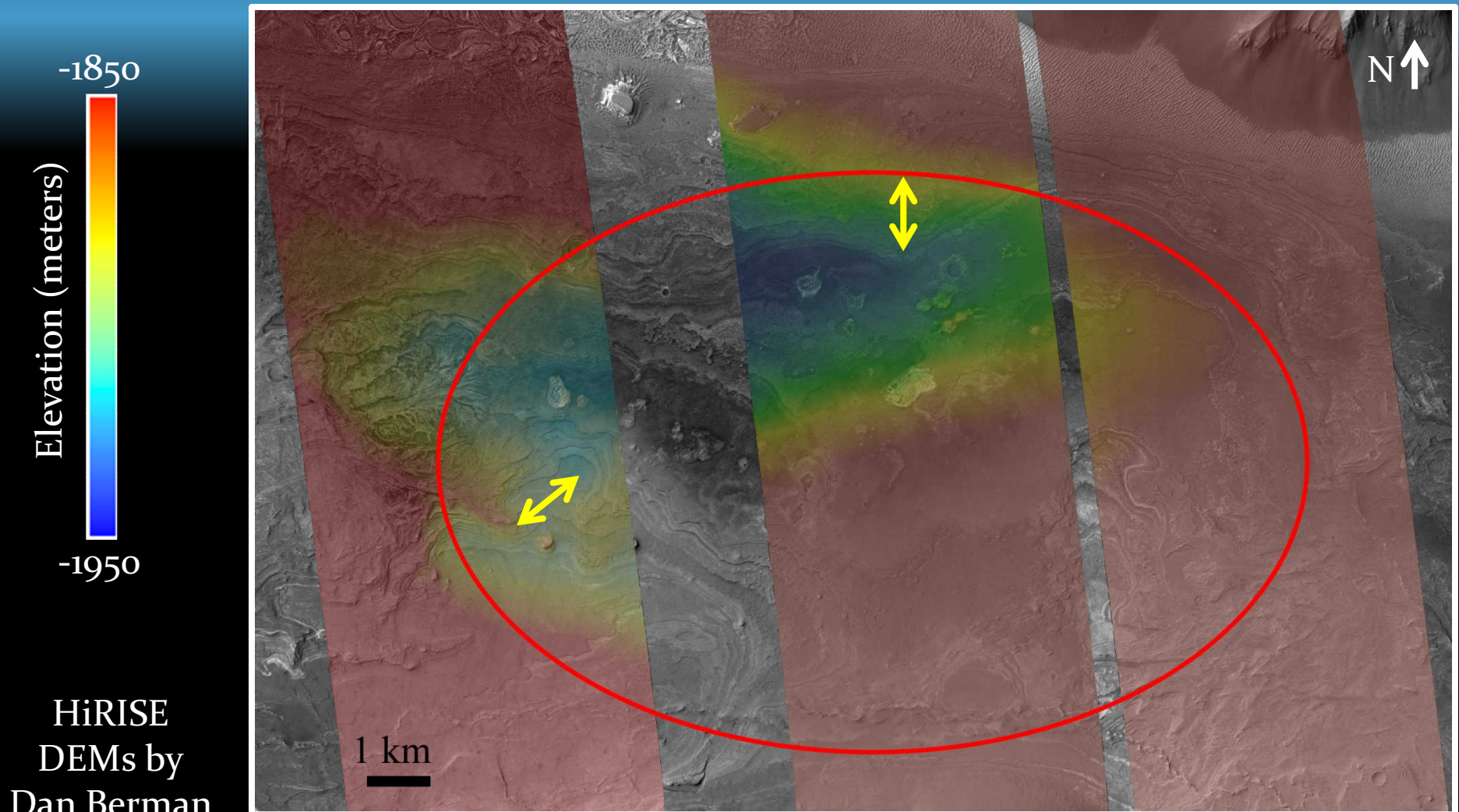
Other Noteworthy Attributes

- Excellent image and spectral dataset for further assessment of landing site
- Preliminary analysis demonstrates that most engineering constraints are met.
 - Detailed route planning needed to ensure rover trafficability over layered terrain.
- Ground-based observations will constrain timing and duration of aqueous history.
 - Ellipse is located on primary target of interest (layered materials)
- Geologic diversity of site enables exploration of a number of scientific questions
 - application to understand similar landforms elsewhere on Mars.
 - How did light-toned layered unit formed?
 - What caused blocky deposits?
 - What is wallrock?
- Spectacular views from a descent imager inside Valles Marineris.

Relief Criteria <100 m relief over 1-1000 m

(proper control authority and fuel consumption during powered descent)

Assessment: few locales with relief in excess of desired range.

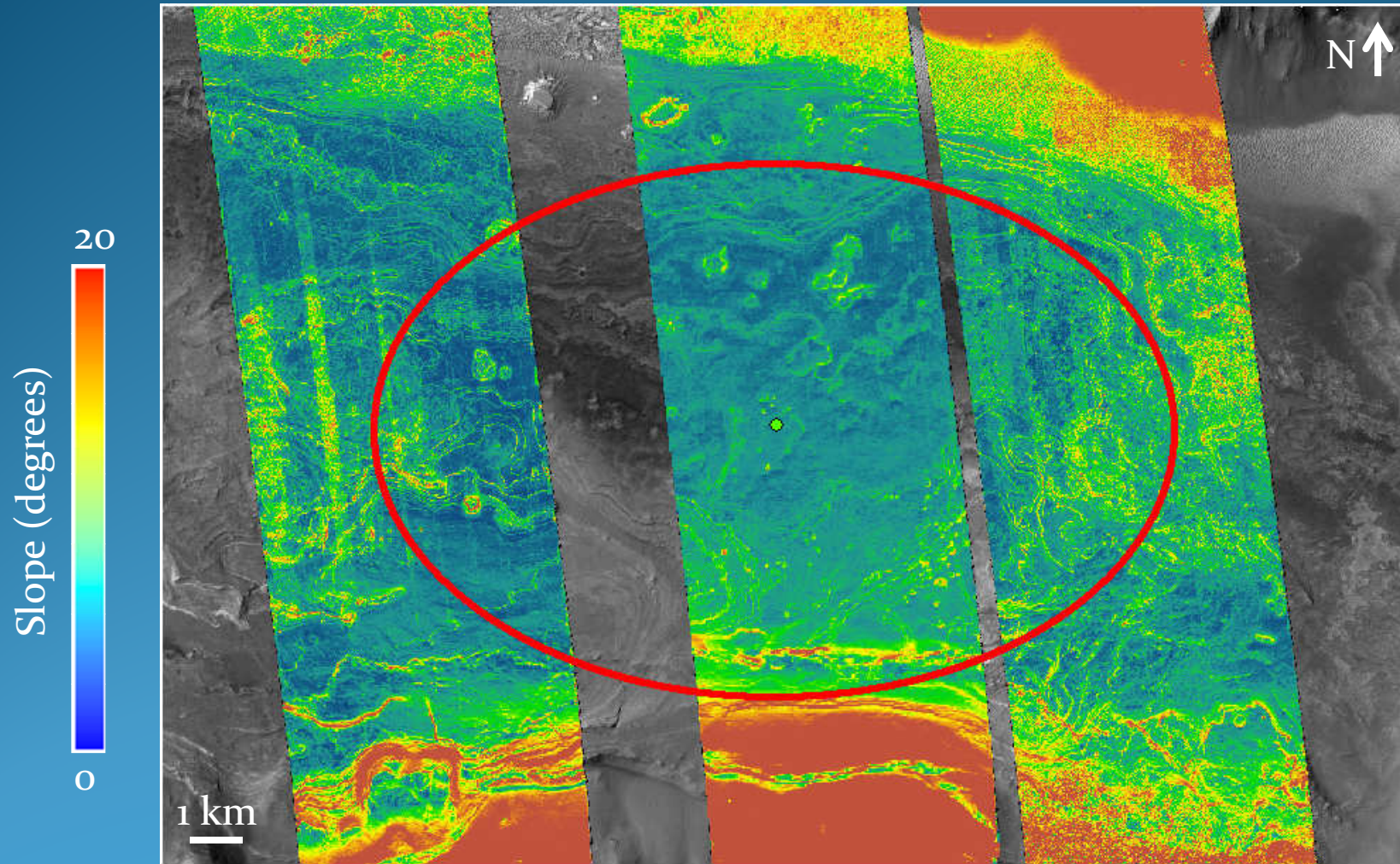


Note: Ellipse is 15 x 10 km

Slope Criteria: $<25^\circ$ over 2-5 m

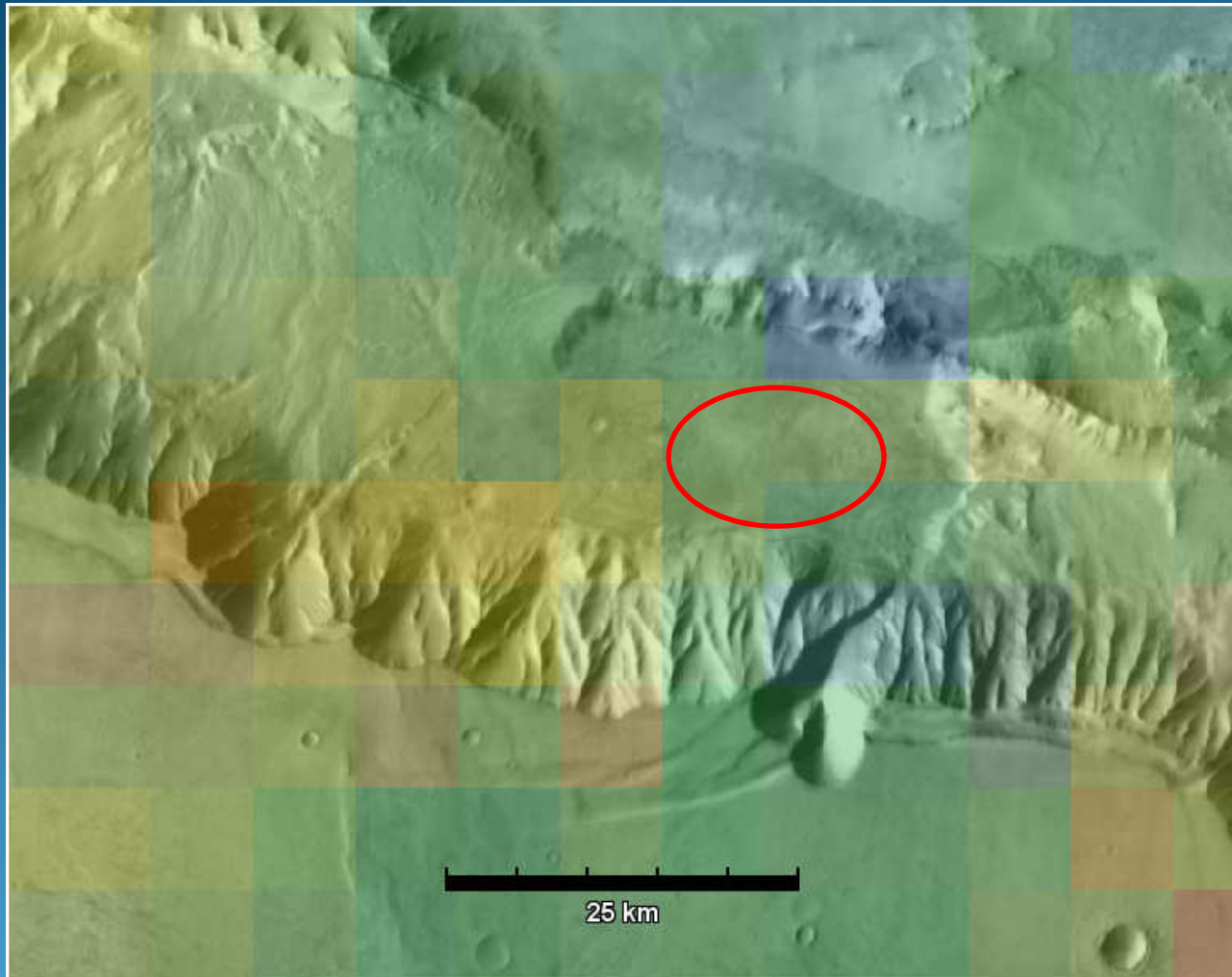
Slope Maps derived from HiRISE DTMs

Assessment: Navigable paths exist, but not all science targets may be accessible.

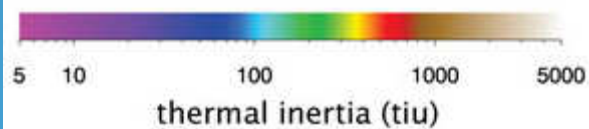


Note: Ellipse is 15 x 10 km

Daytime Thermal Inertia

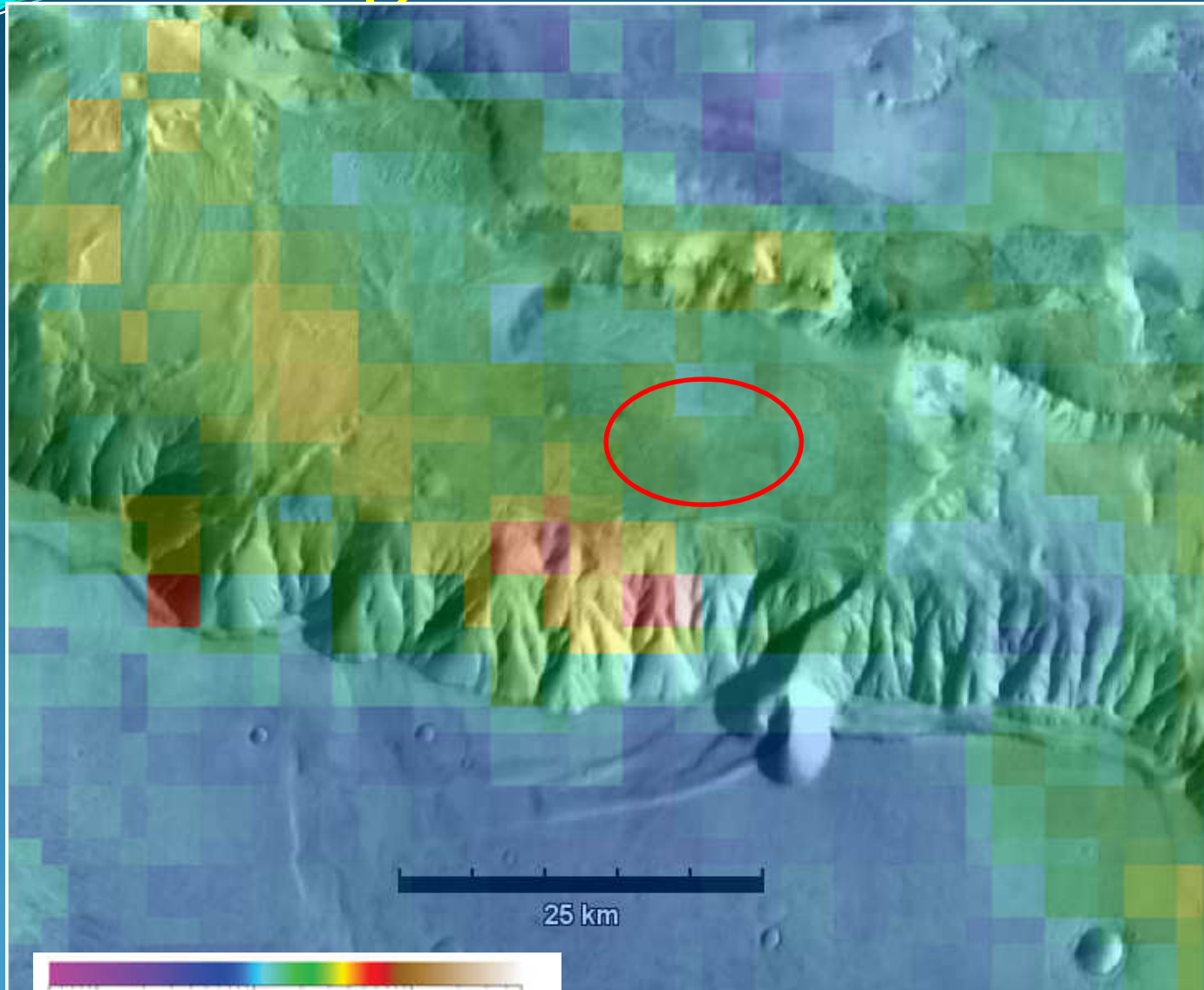


Meets criteria.

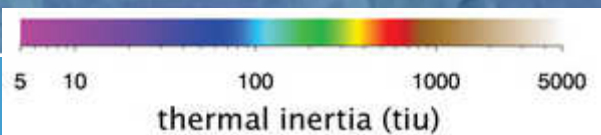


Note: Ellipse is 15 x 10 km
(Putzig and Mellon, 2007)

Nighttime Thermal Inertia



Meets criteria.



Note: Ellipse is 15 x 10 km
(Putzig and Mellon, 2007)