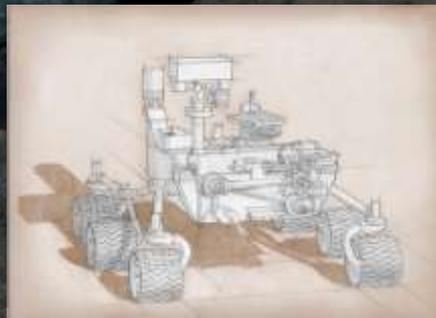


Mawrth Vallis clay unit: probing the early Mars Habitability, Climate and Origin of Life

*D. Loizeau, F. Poulet, B. Horgan,
N. Mangold, J. Michalski, J. Bishop*

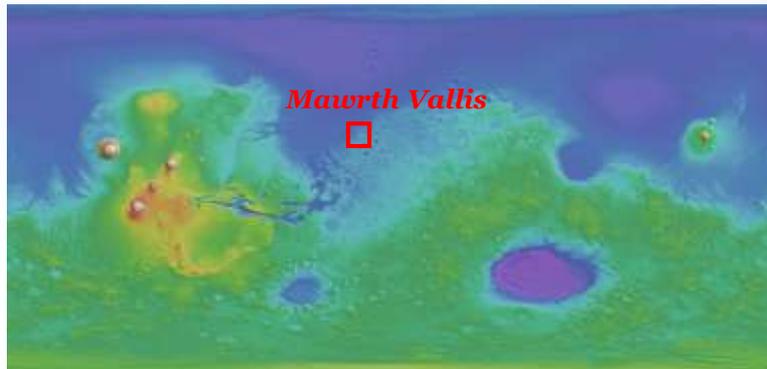
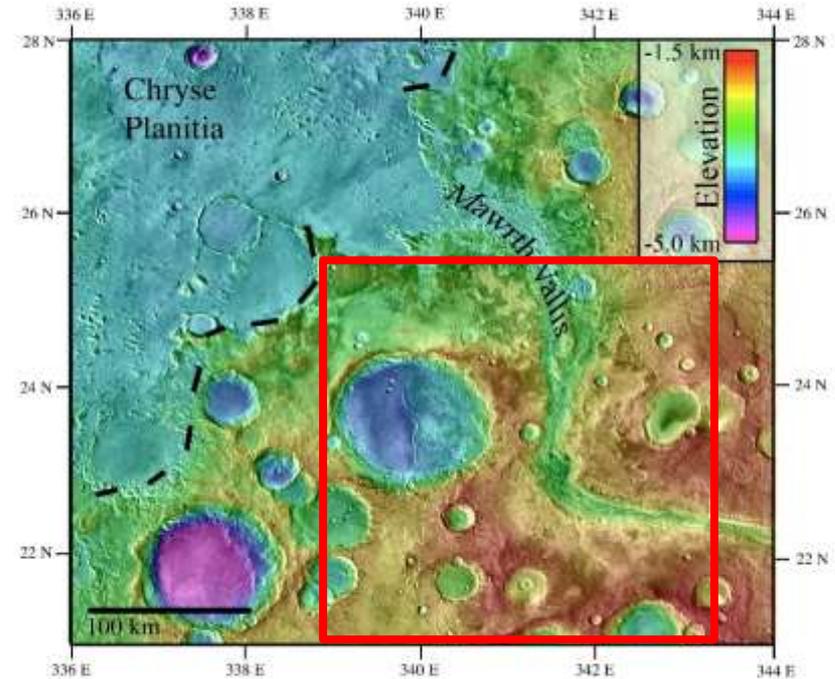
Mars 2020 Landing Site Workshop

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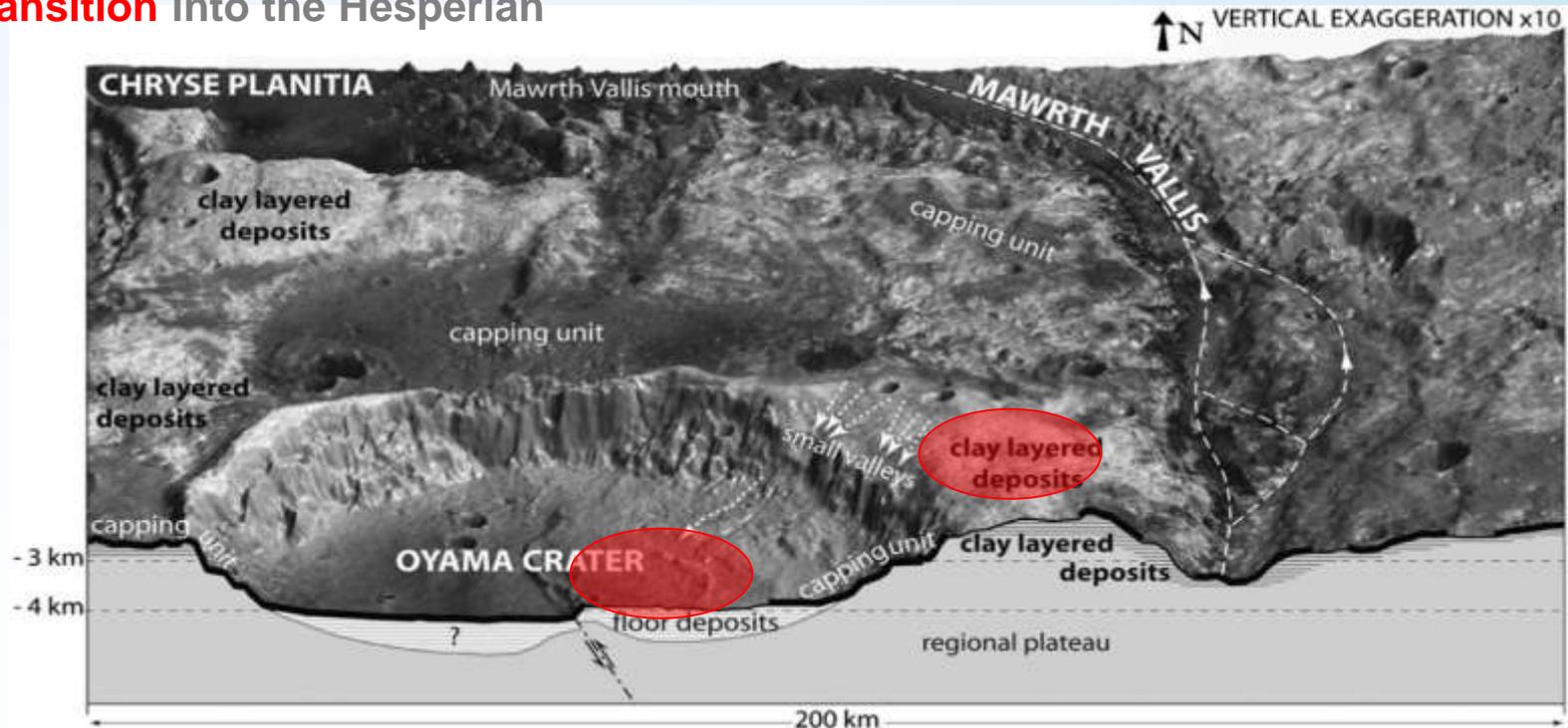
Regional Context

- Extended unit with a large variety of hydrated minerals associated with layers (several ellipses were studied)
- Largest clay content identified on Mars
- Numerous studies (40+ peer-reviewed papers including science and nature papers)



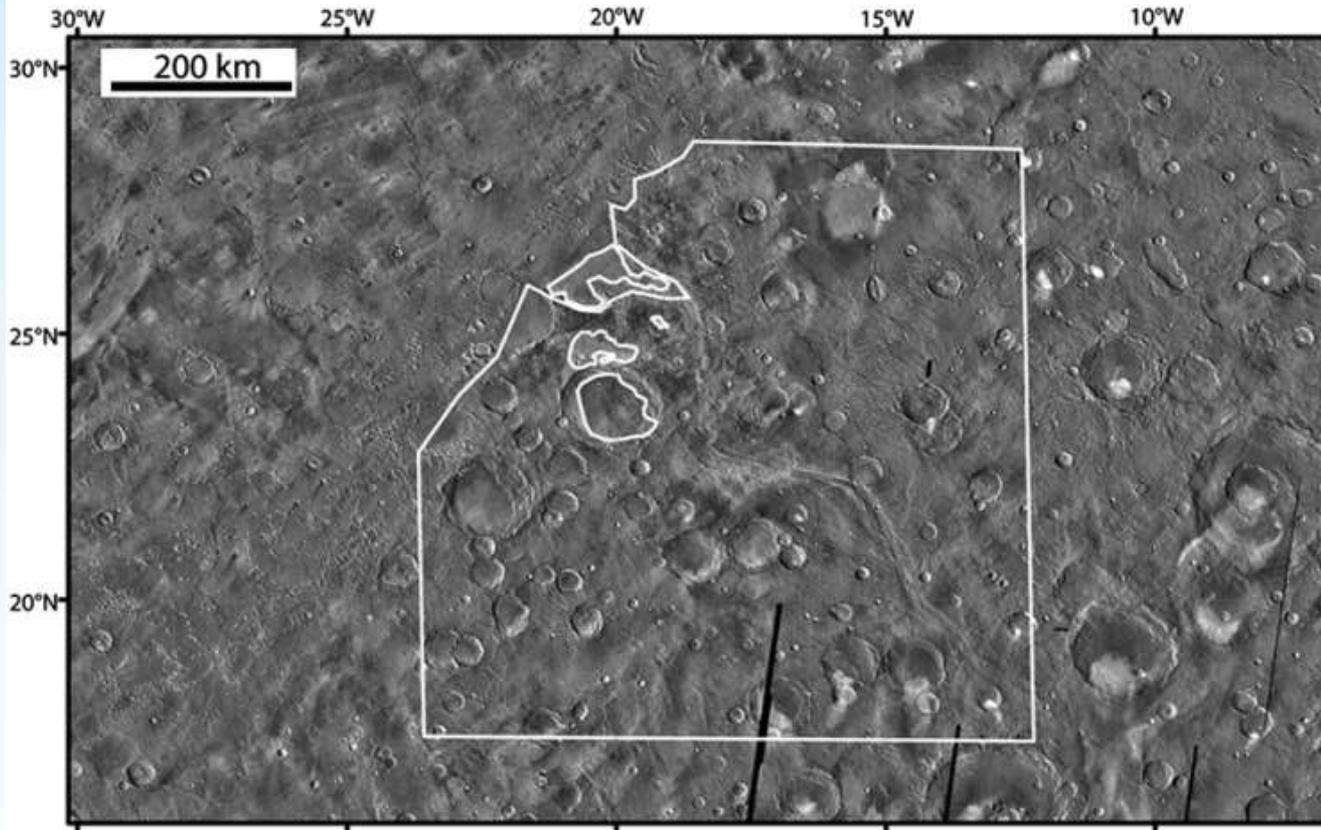
Attractive points

1. **Mineralogically very** diverse site
2. **Lithologically** diverse site that captures multiple environments
3. Both **ancient altered Noachian deposits** and **remobilized sediments**
4. Extremely **ancient section** of rocks probing an enigmatic epoch in Solar System history
5. **Several** types of science (**astrobiological**) targets
6. Opportunity to **sample rocks** from the deep Noachian up **through the global transition** into the Hesperian



Early Noachian to the Hesperian

MARS 2020 at MAWRTH VALLIS



model age: ~4.0 Ga ~3.9 Ga ~3.8 Ga ~3.7 Ga-3.6 Ga

**MAWRTH VALLIS
UNITS**

regional plateau inter-crater plateaus Oyama crater floor deposits dark capping unit & Chryse Planitia filling

— PROBABLE AQUEOUS ALTERATION PERIOD —

corresponding process:

REGIONAL PLATEAU
LARGELY FORMED

LAYERED UNIT
LARGELY DEPOSITED

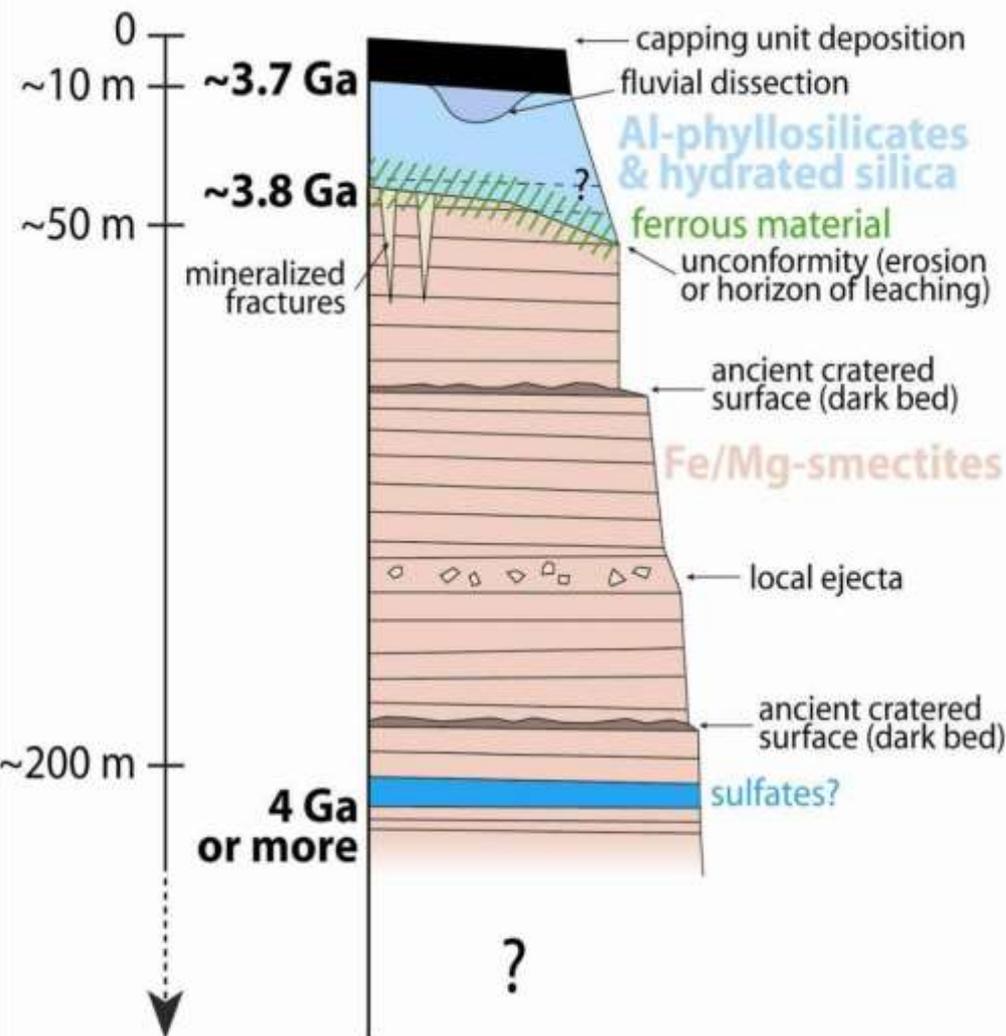
EROSION,
REDEPOSITION,
FLUVIAL ACTIVITY

NO MORE
SURFACE
ALTERATION

WIND
EROSION

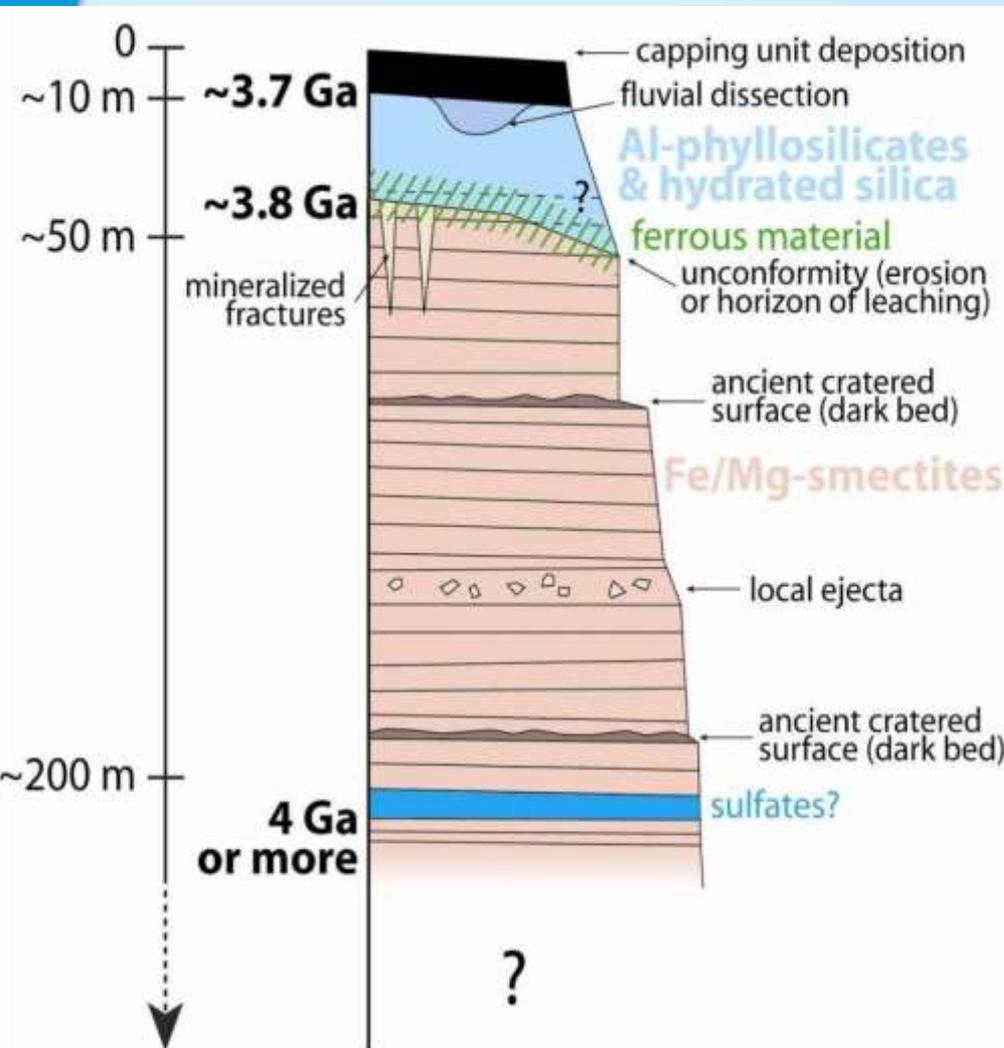
Epochs: Early Noachian | Middle Noachian | Late Noachian | Early to Late Hesperian ... Amazonian

Complex aqueous diversity on Mars



Gale Crater	Mawrth Vallis
Fe/Mg-phyllsilicate	Fe/Mg-phyllsilicate
	Al-phyllsilicate
	Hydrated silica
	Kaolinite
	Acid-treated clays
Sulfates (polyhydrated)	Sulfates (bassanite jarosite)
	Alunite

Complex aqueous history on Mars



Phase 1

Progressive deposition and alteration of sediments or pyroclastic deposits => smectites

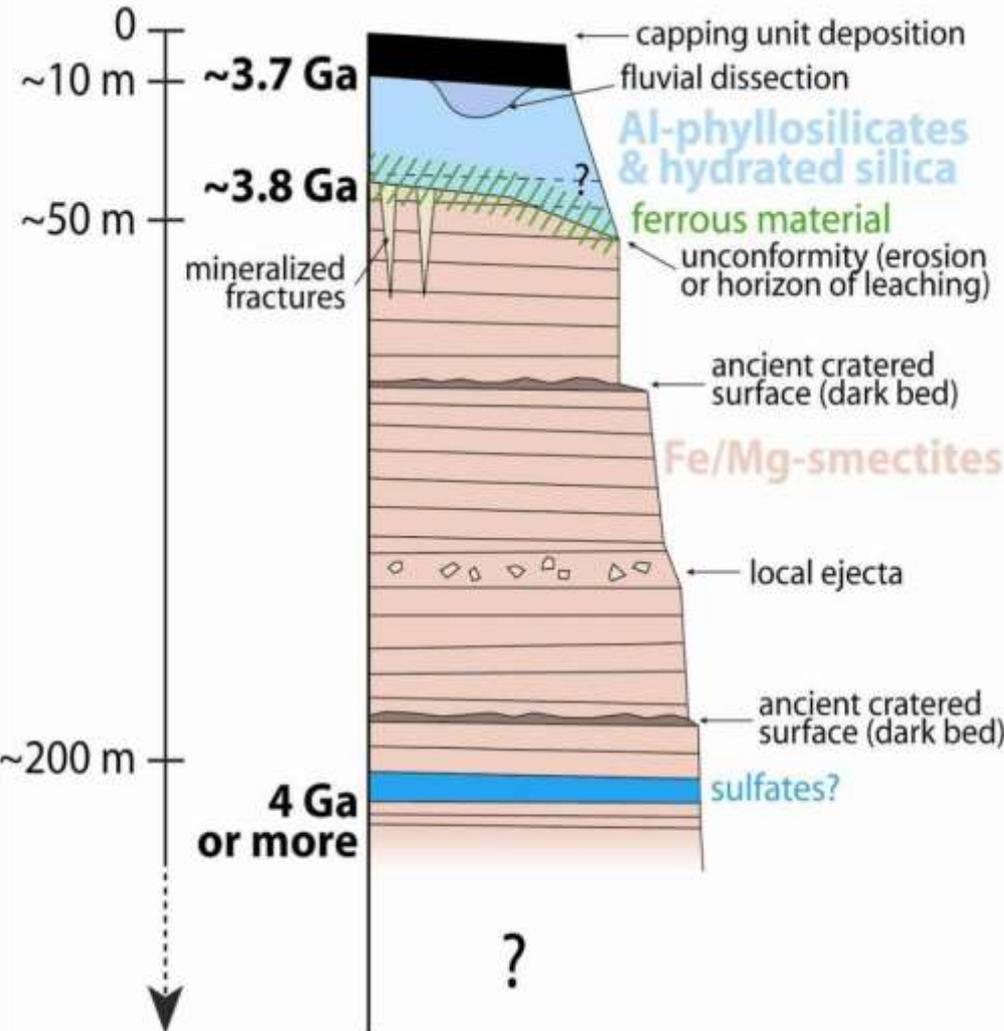
Early to middle Noachian

Moderate water rock ratio

Local precipitation of sulfates

Complex aqueous history on Mars

Phase 2



Continued surface weathering, greater or acidic leaching of the surface layers

Surface assemblage: Kaolins, alunite, ferrous clays

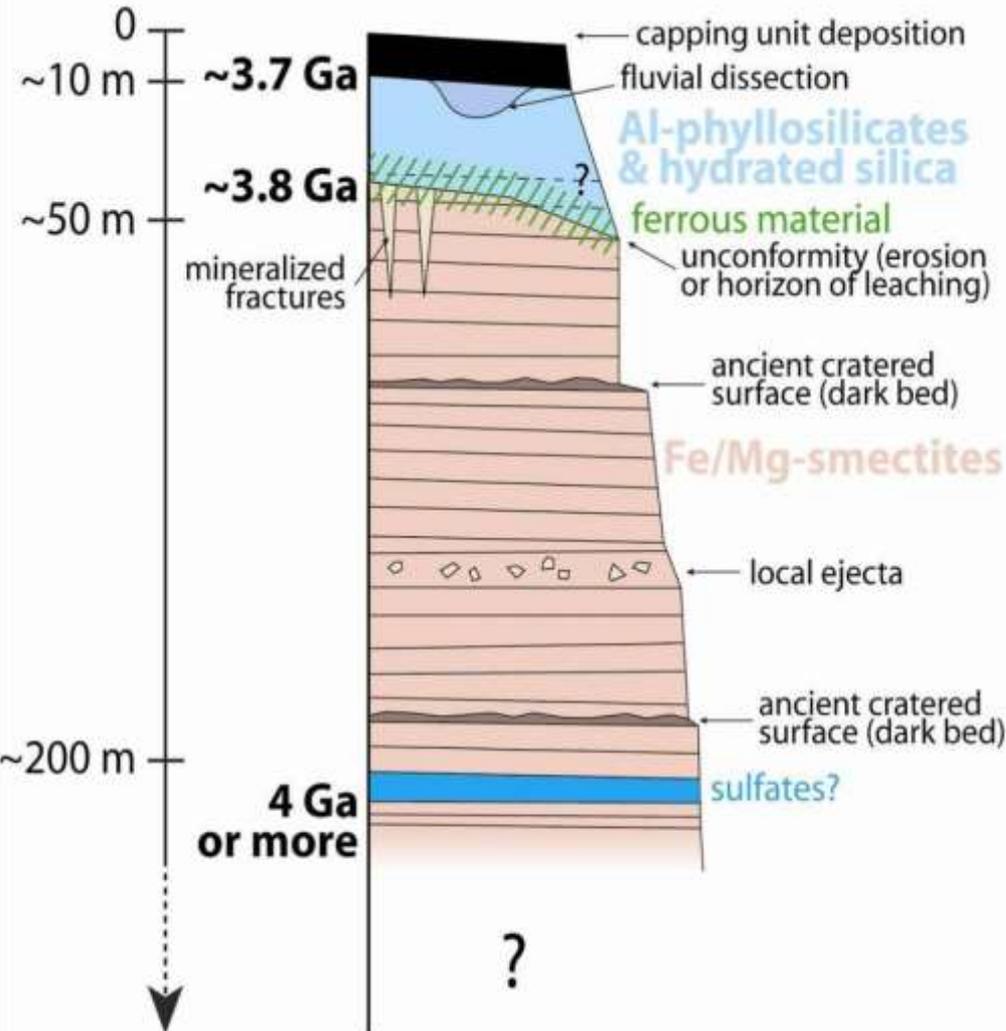
Local precipitation of sulfates (jarosite)

Extreme Fe/S redox gradient: excellent **source of energy** for likely Martian microbes.

Fluid circulation in fractures => halo-bounded fractures

Complex aqueous history on Mars

Phase 3



Pyroxene-bearing dark cap deposition

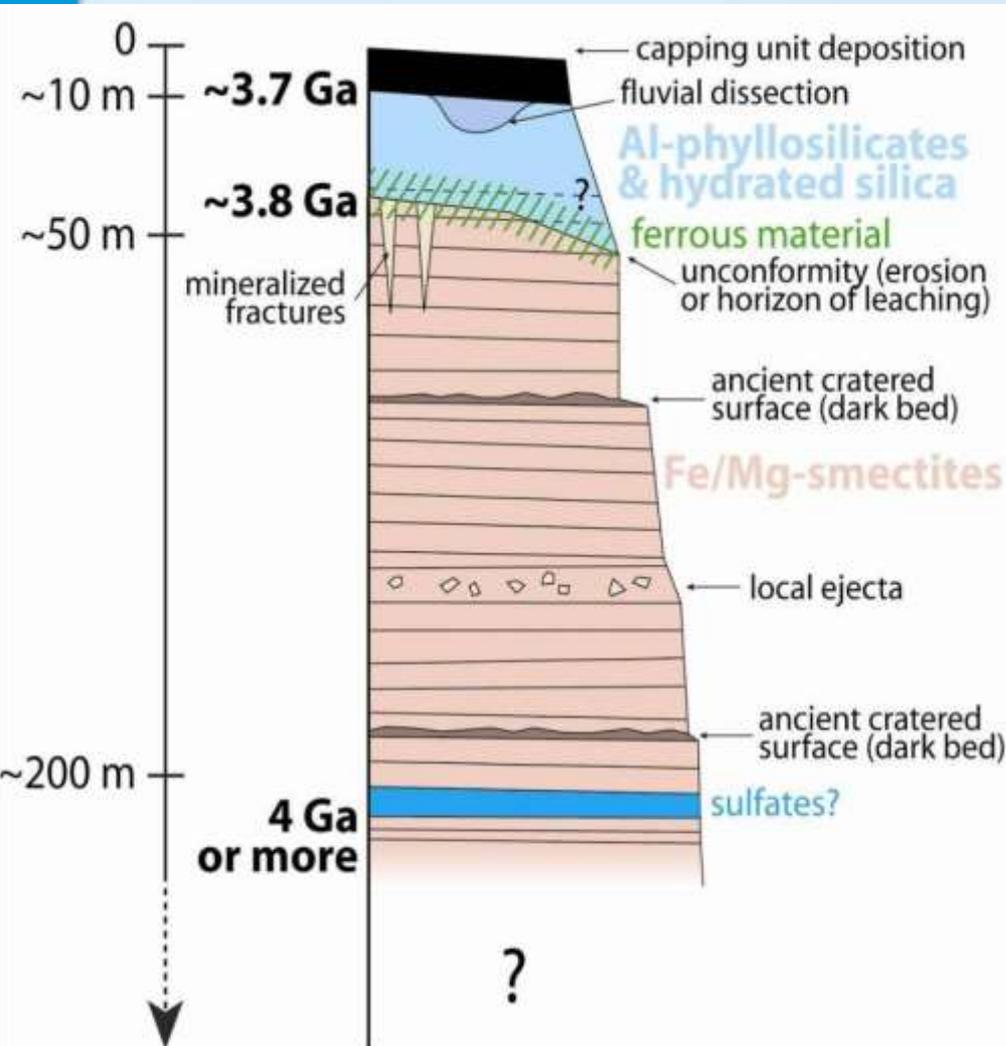
Early Hesperian

Probable volcanic/pyroclastic deposits

No more aqueous alteration

Preservation of clays and morphologic features (inverted valleys)

Complex aqueous history on Mars



Phase 4

Wind erosion

the whole section is progressively and continuously exhumed

Hesperian and Amazonian

Best Terrestrial Analogs

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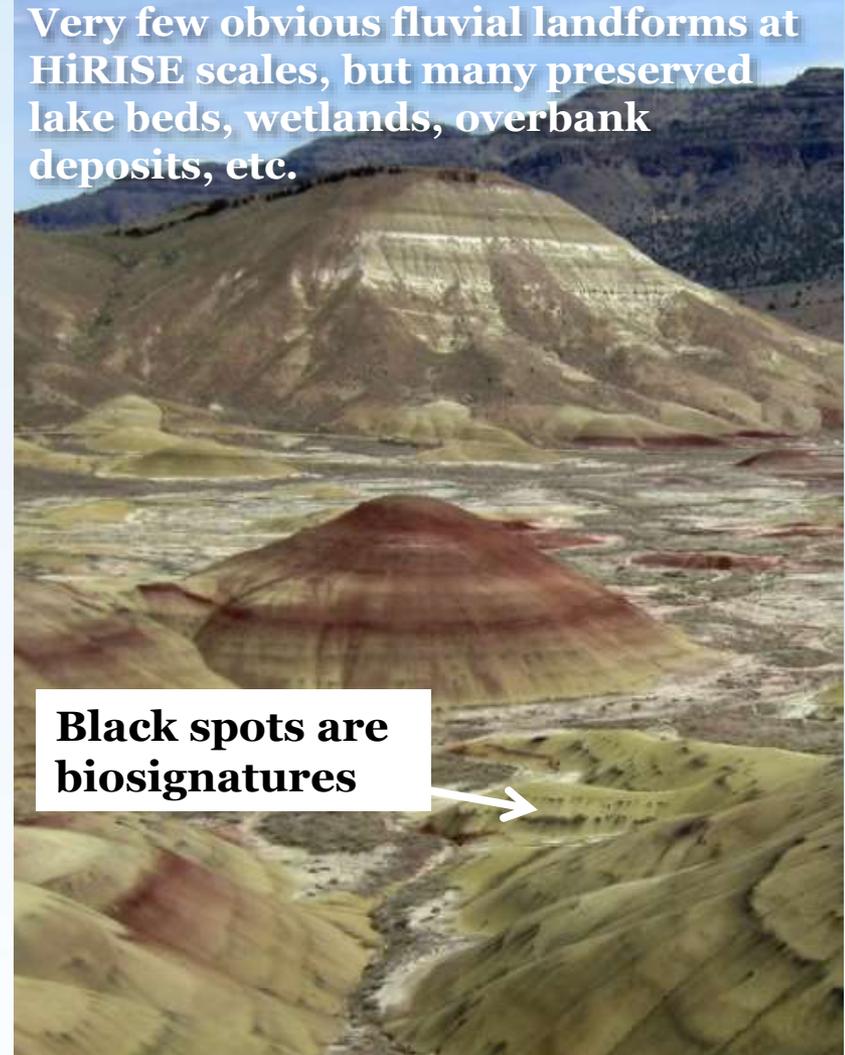
Painted Desert, AZ: Floodplain soils formed in fluvial/ lacustrine/deltaic/aeolian sediments (~200 My)

Some fluvial landforms at HiRISE scales, but many other aqueous environments as well



John Day Fossil Beds, OR: Soils formed in pyroclastic sediments, modified by local fluvial/lacustrine activity (~30 My)

Very few obvious fluvial landforms at HiRISE scales, but many preserved lake beds, wetlands, overbank deposits, etc.



Black spots are biosignatures



Habitability and Preservation of Organics

- Soils are **highly habitable** environments - energy, water, nutrients, etc.
- High clay content and rapid burial leads to **organic preservation** in paleosols
- **Concentration** of organics is overall low-mod, but can be **locally high**
- **Reducing soils** cause immediate **preservation** and can lead to concentrated organics in wetlands
- **Reduced paleosols at Mawrth** shall be excellent targets for *in situ* search for **organics and biosignatures**



Archean paleosols preserve the oldest non-marine organics on Earth (1-3 Gy)

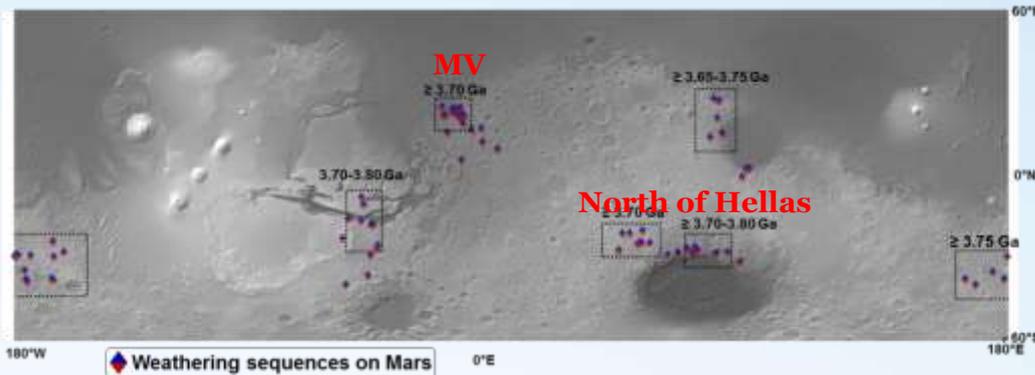


Dorset, UK tonstein. Mineralogical analog for Mawrth Vallis. Organic-rich!

Probing for past climatic conditions

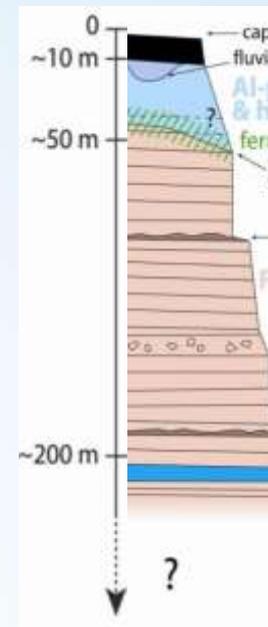
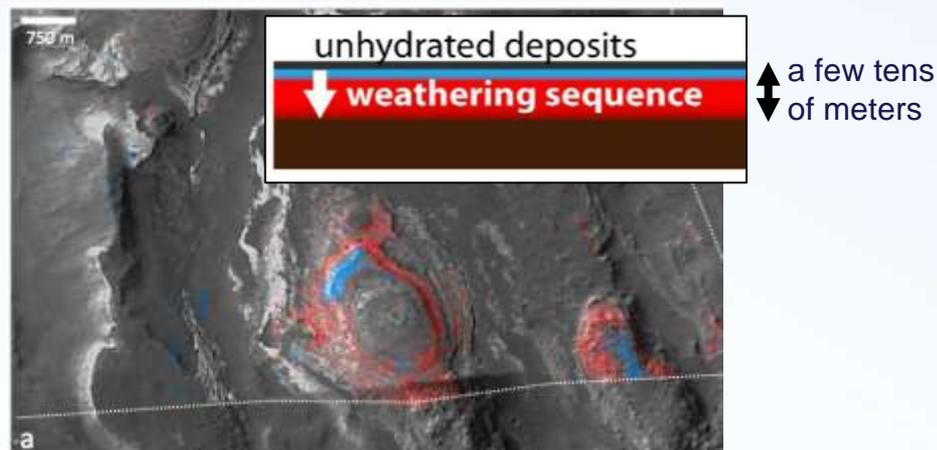
Al-clays over Fe/Mg-smectites sequence is common on Mars during this time period:

⇒ MV and its weathering profile allow to characterize Mars' ancient climate and underlying processes



Mawrth sequence is several time thicker than others and mineralogy much more complex

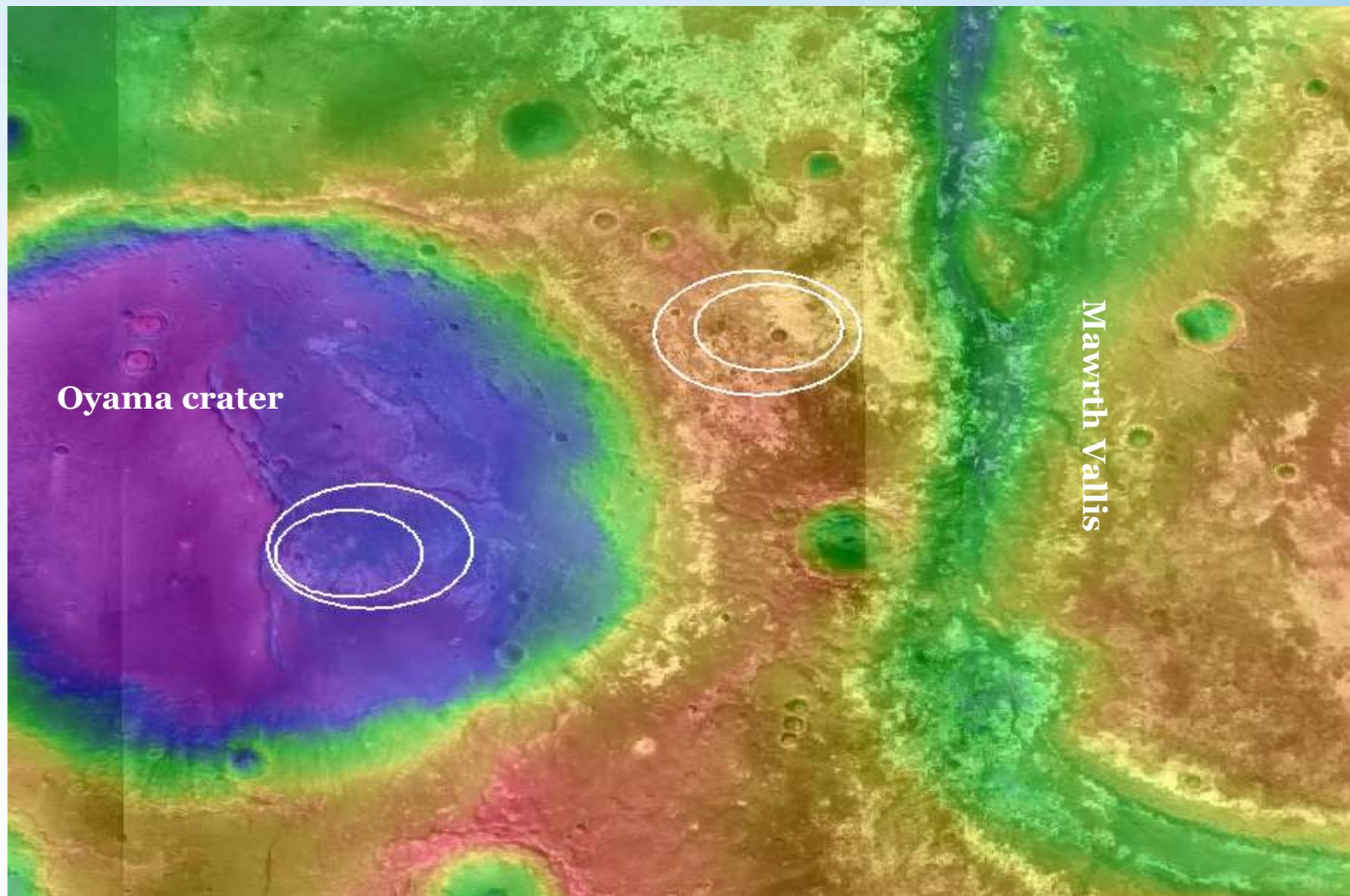
Consistent with long-term (~million years) leaching profiles in a wetter climate



Carter et al. (2015)

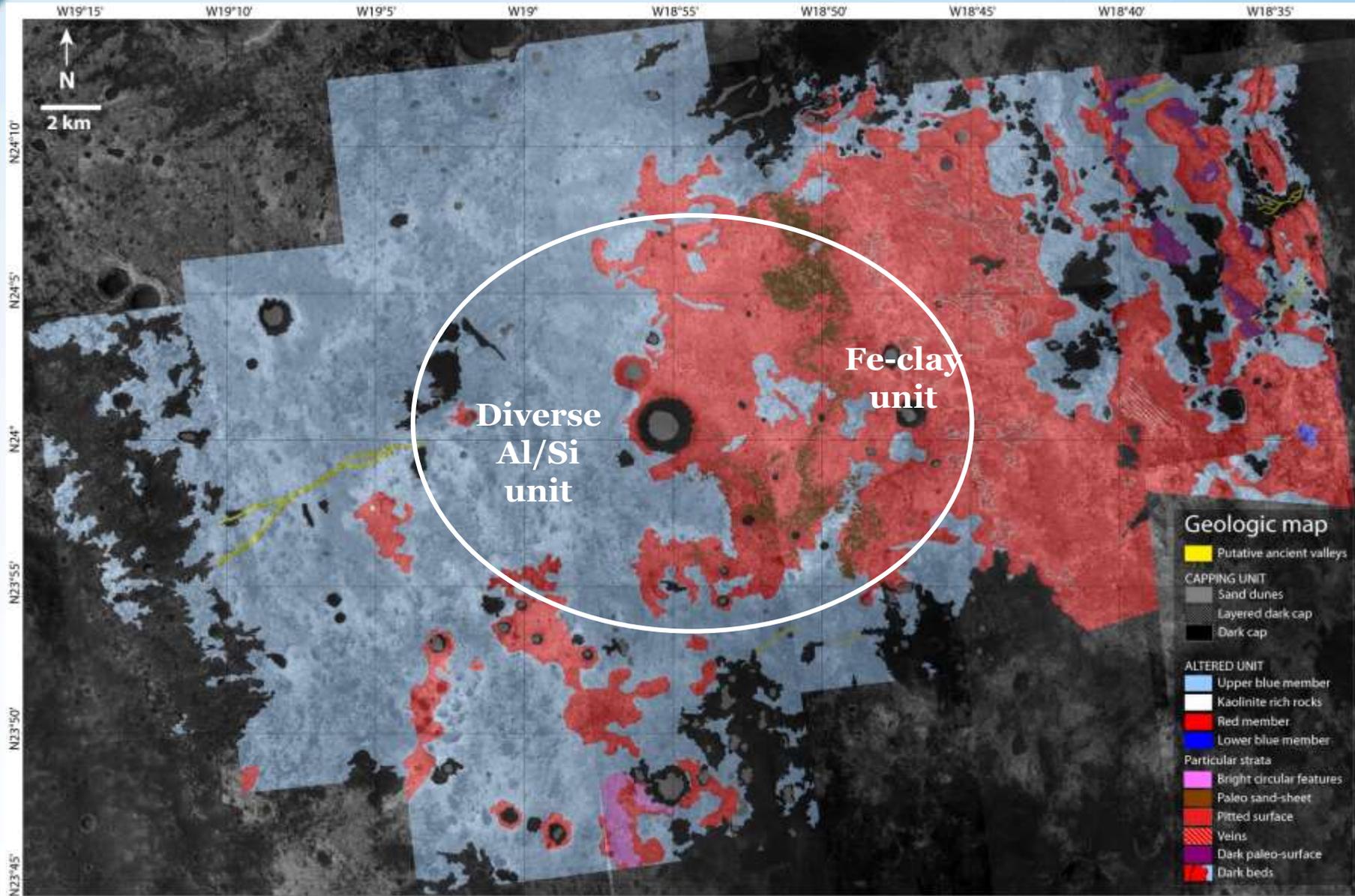
Two ellipses: *Plateau ellipse* and *Oyama*

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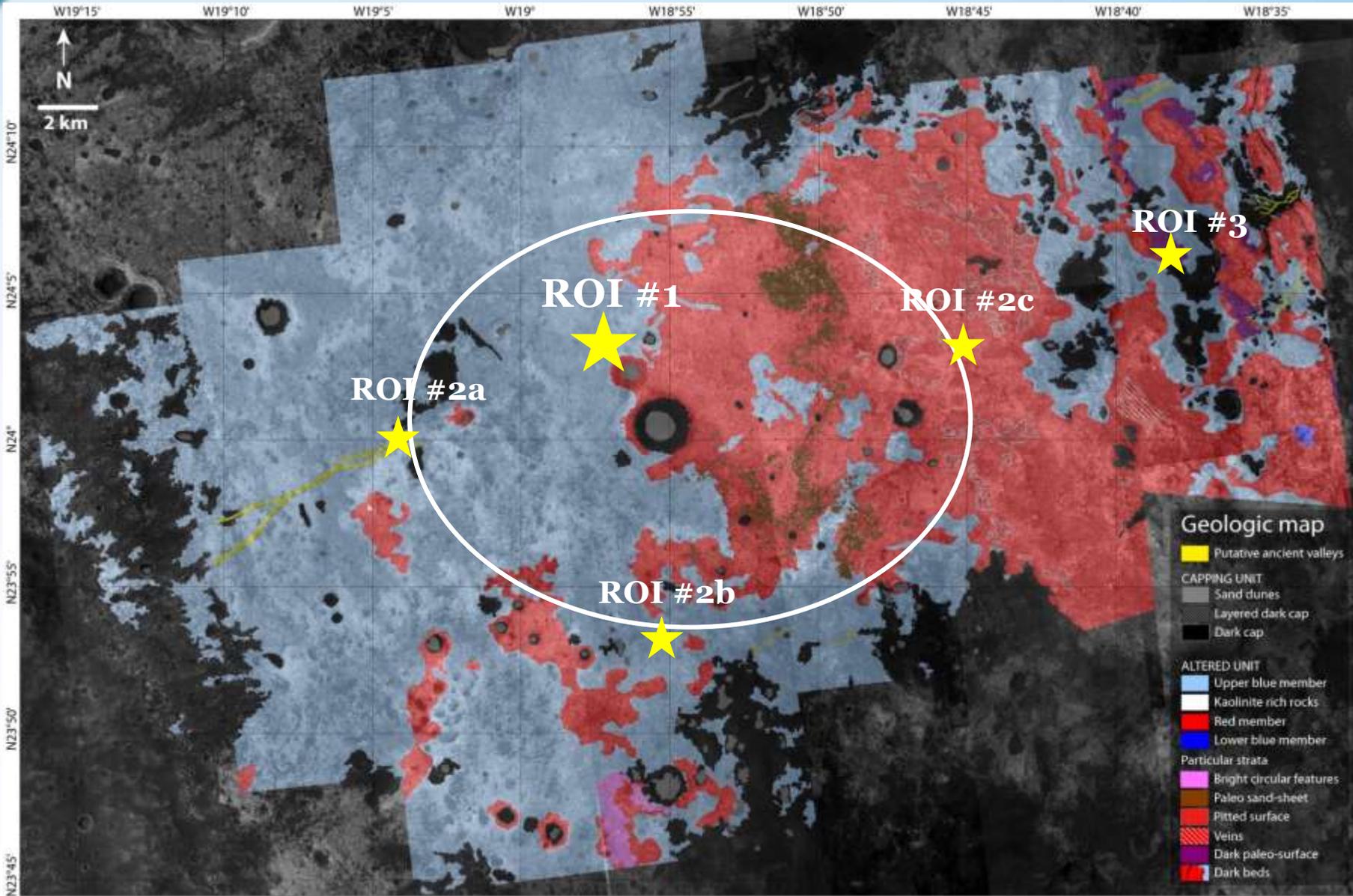


GEOLOGIC MAP

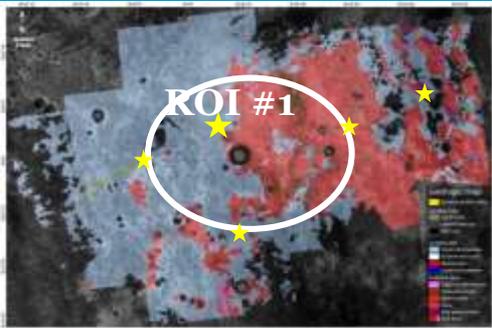
MARS 2020 at MAWRTH VALLIS



ROIs

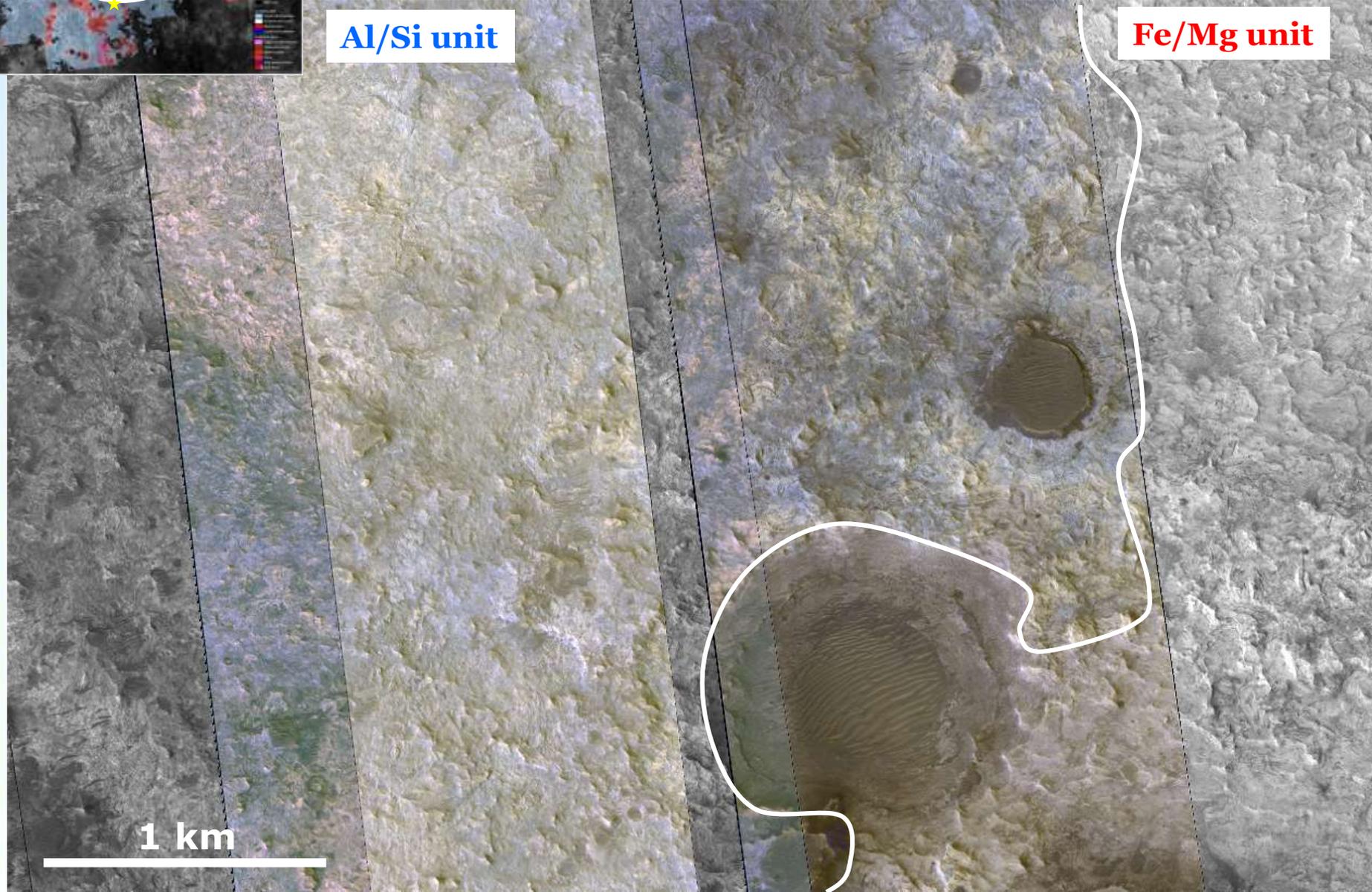


ROI 1



Al/Si unit

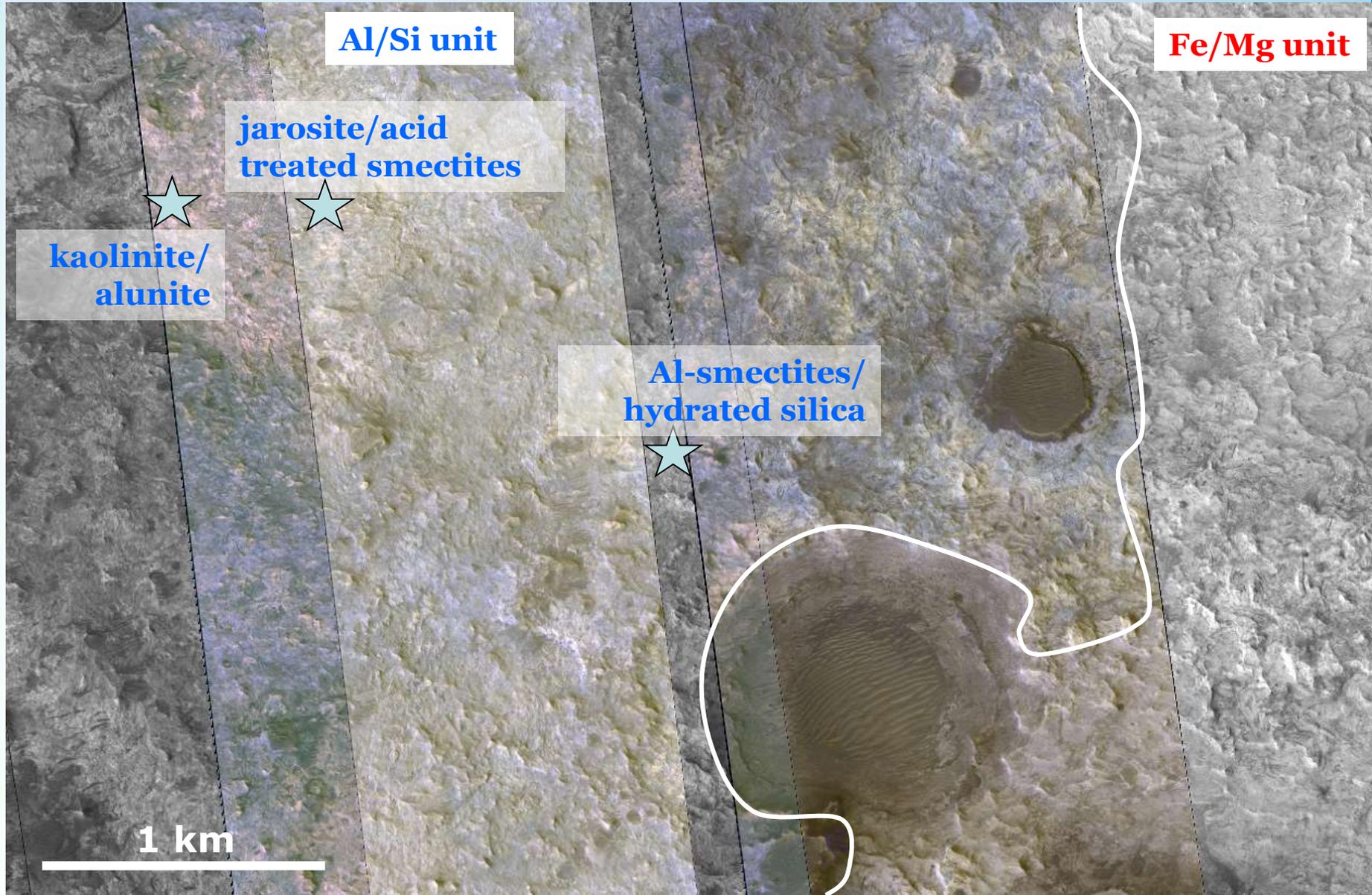
Fe/Mg unit



1 km

ROI 1

MARS 2020 at MAWRTH VALLIS



Al/Si unit

Fe/Mg unit

**jarosite/acid
treated smectites**

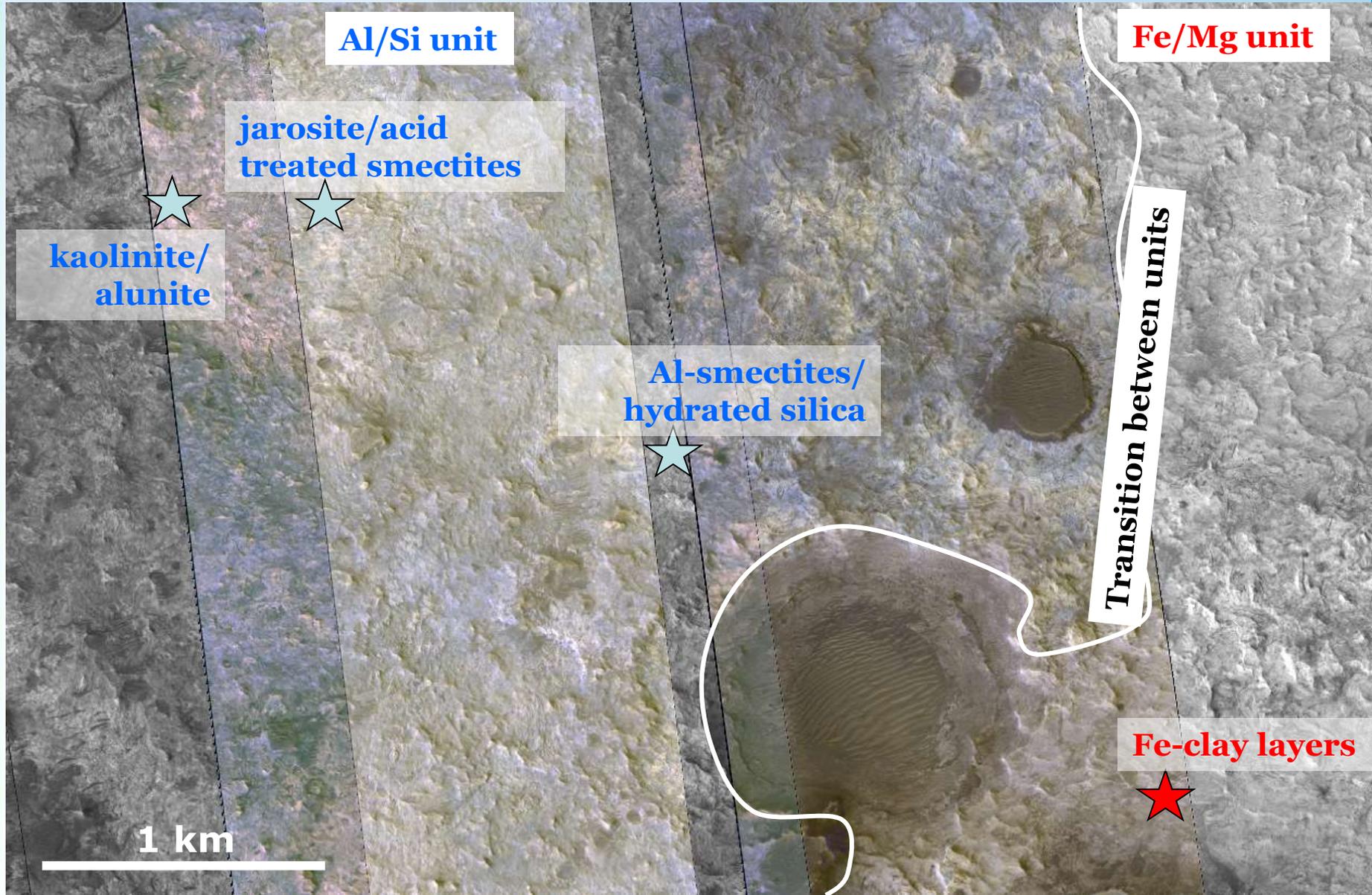
**kaolinite/
alunite**

**Al-smectites/
hydrated silica**

1 km

ROI 1

MARS 2020 at MAWRTH VALLIS



Al/Si unit

Fe/Mg unit

jarosite/acid
treated smectites

kaolinite/
alunite

Al-smectites/
hydrated silica

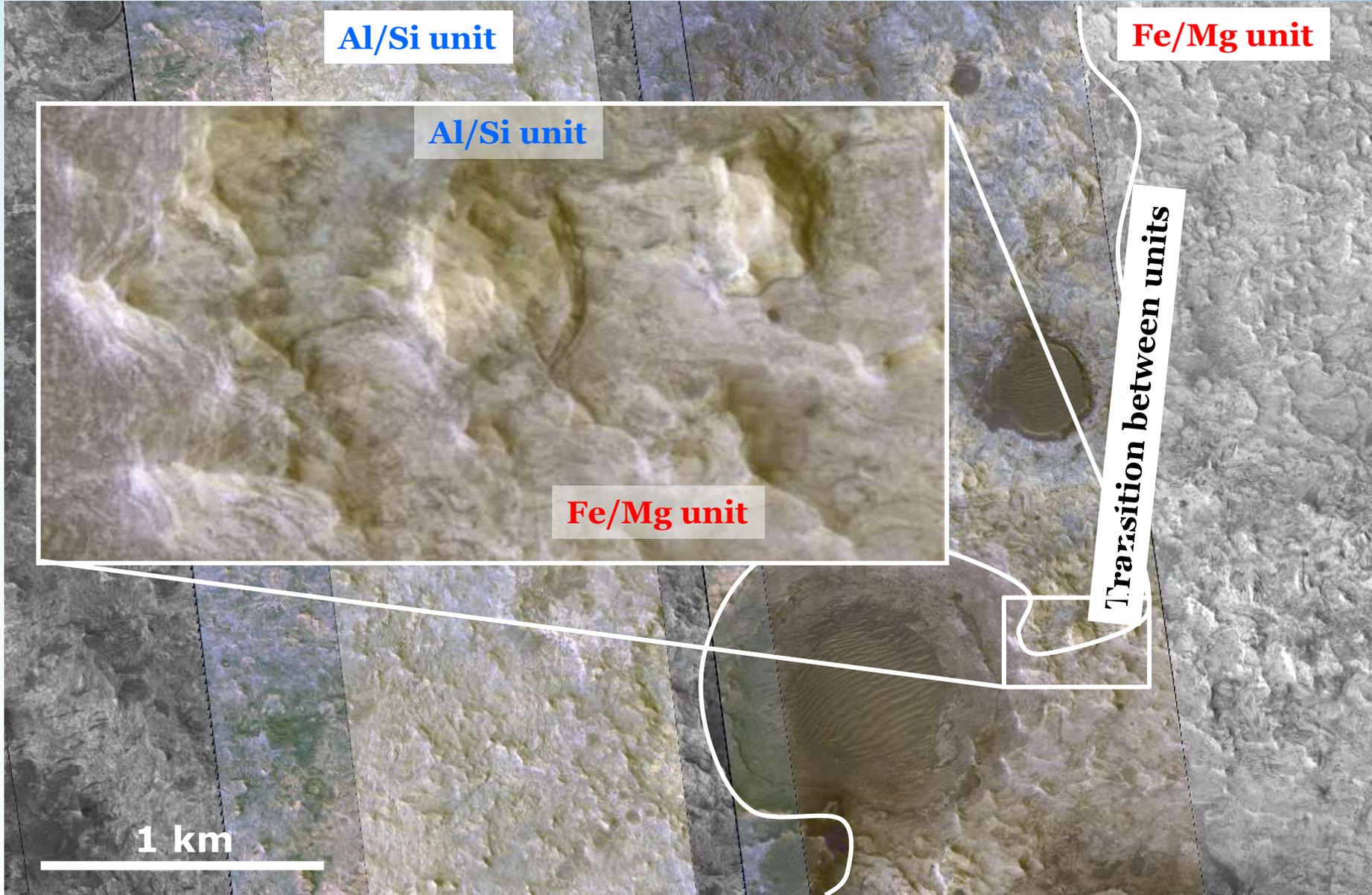
Transition between units

Fe-clay layers

1 km

ROI 1

MARS 2020 at MAWRTH VALLIS



Al/Si unit

Fe/Mg unit

Al/Si unit

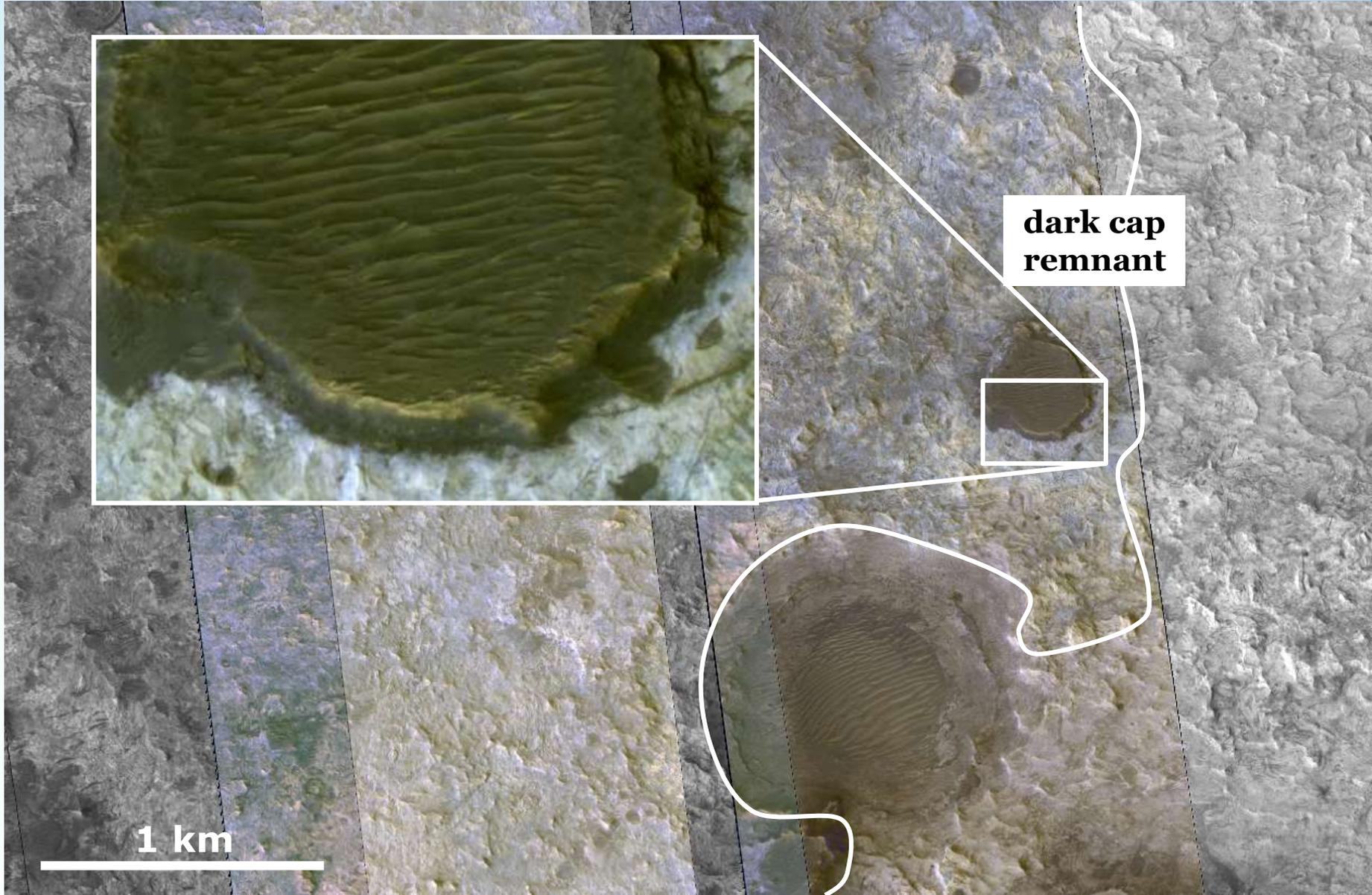
Fe/Mg unit

Transition between units

1 km

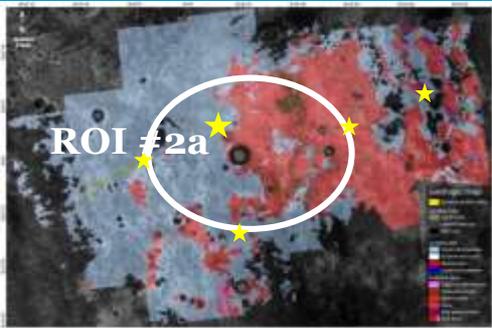
ROI 1

MARS 2020 at MAWRTH VALLIS

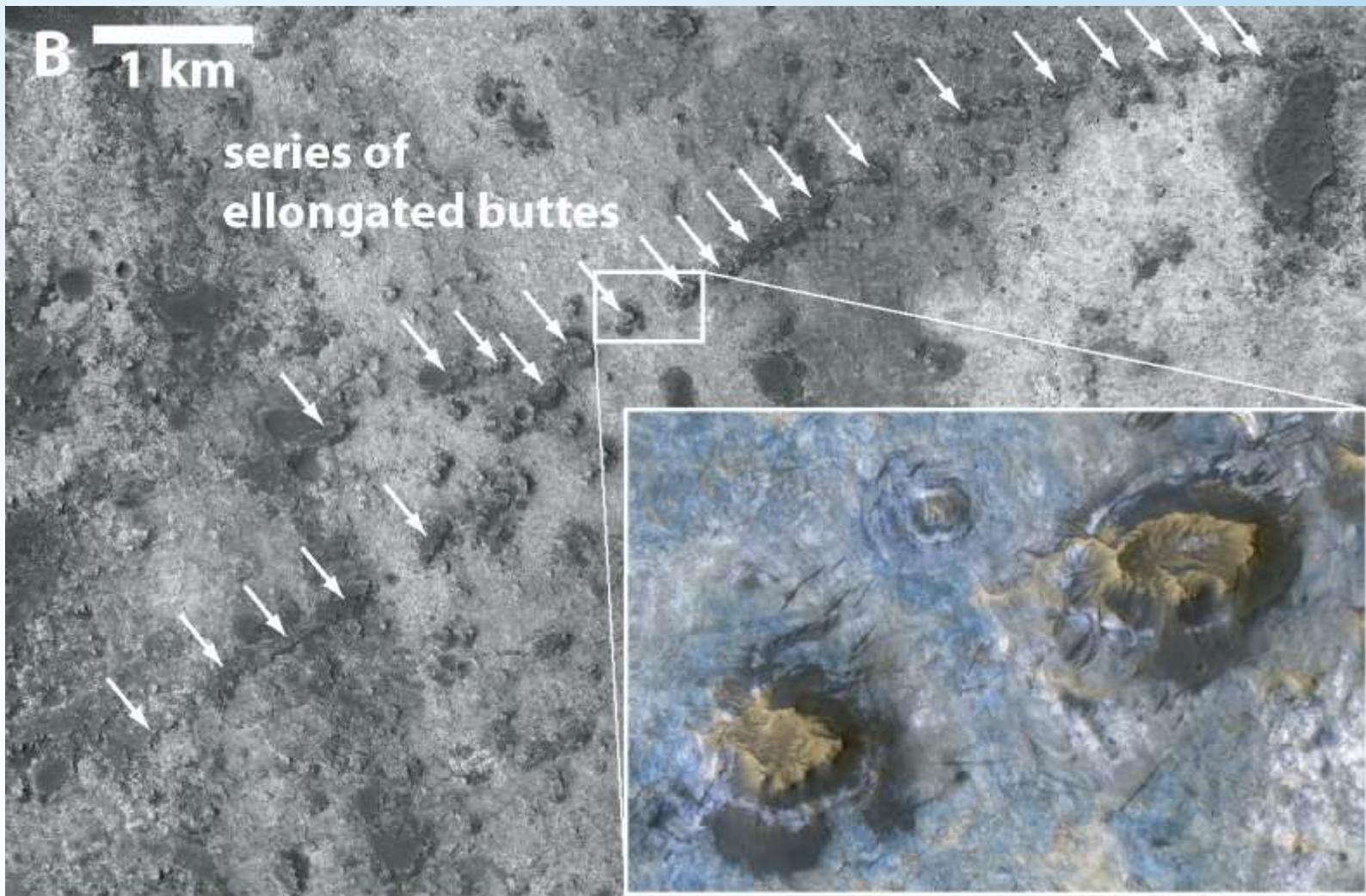


1 km

dark cap
remnant

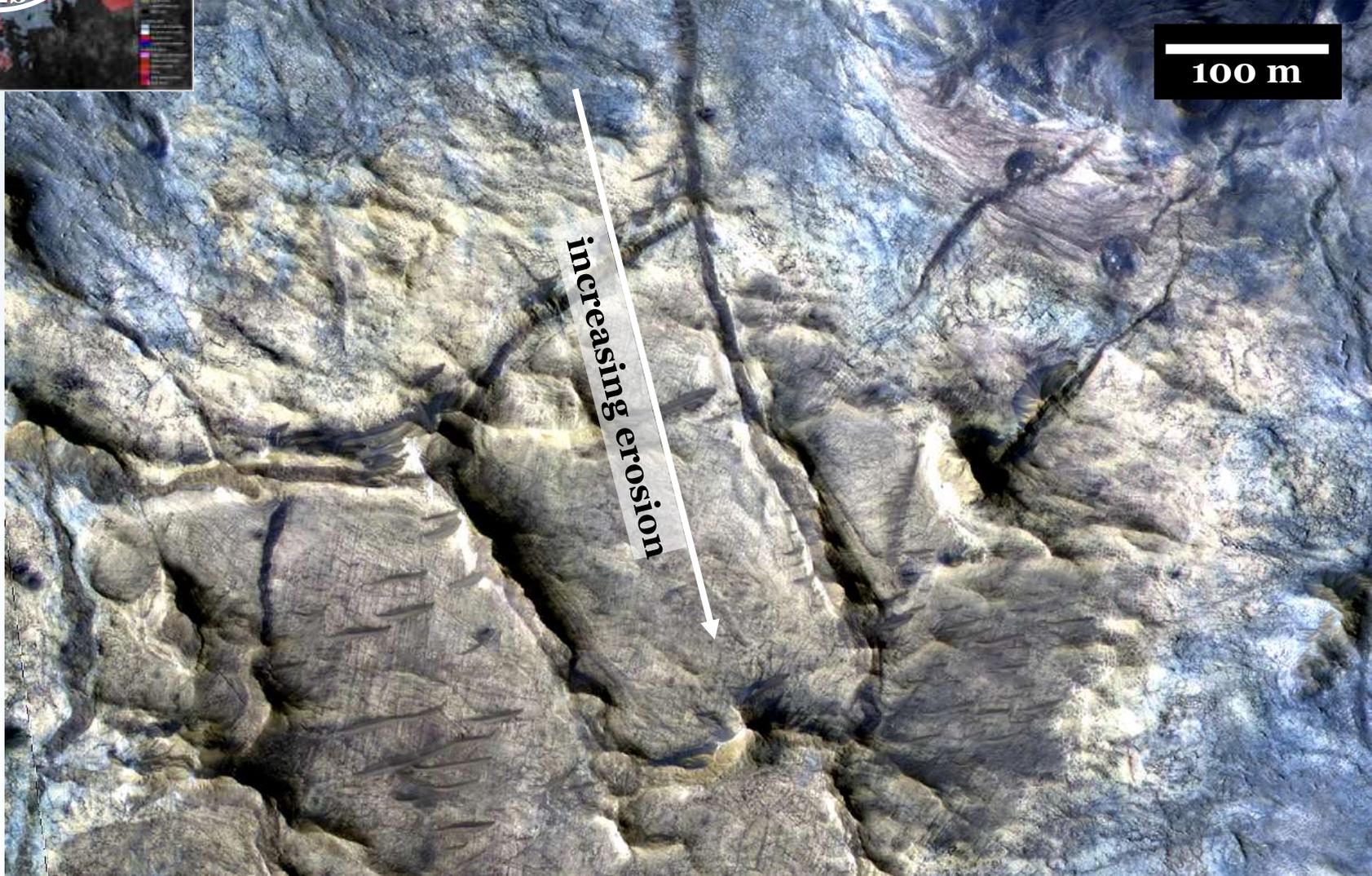
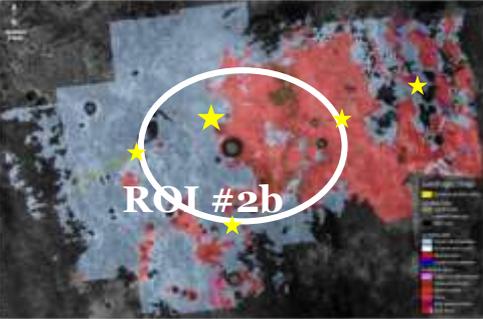


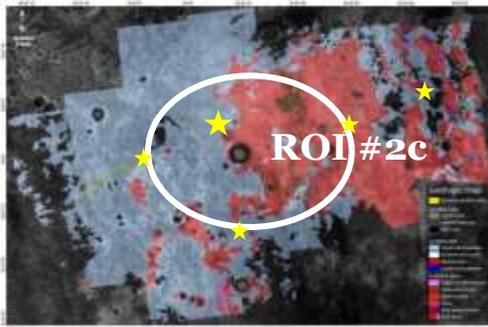
ROI: ancient inverted valleys



ROI: Halo-bounded fractures and veins

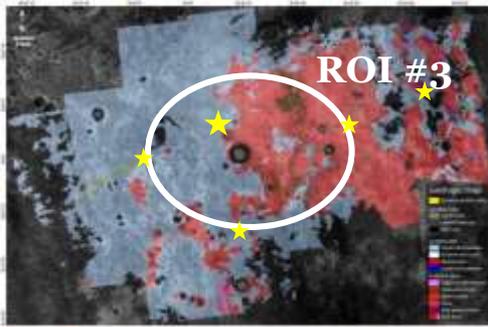
MARS 2020 at MAWRTH V



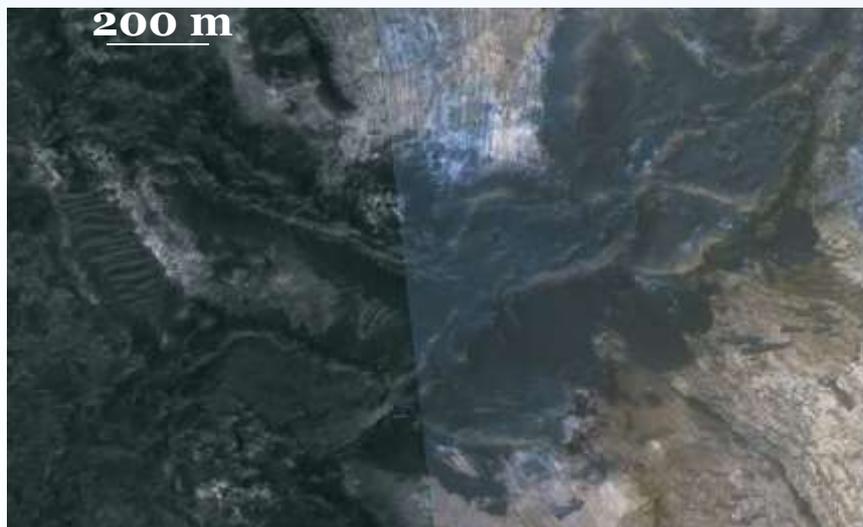
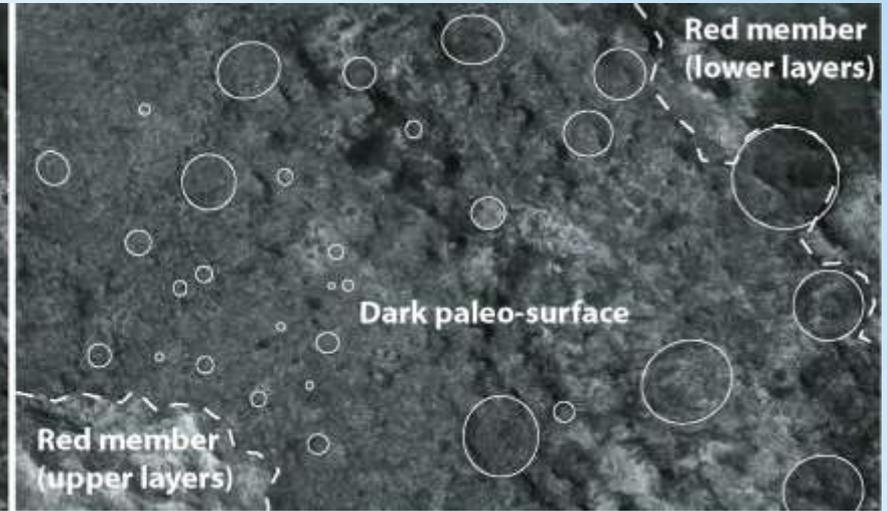
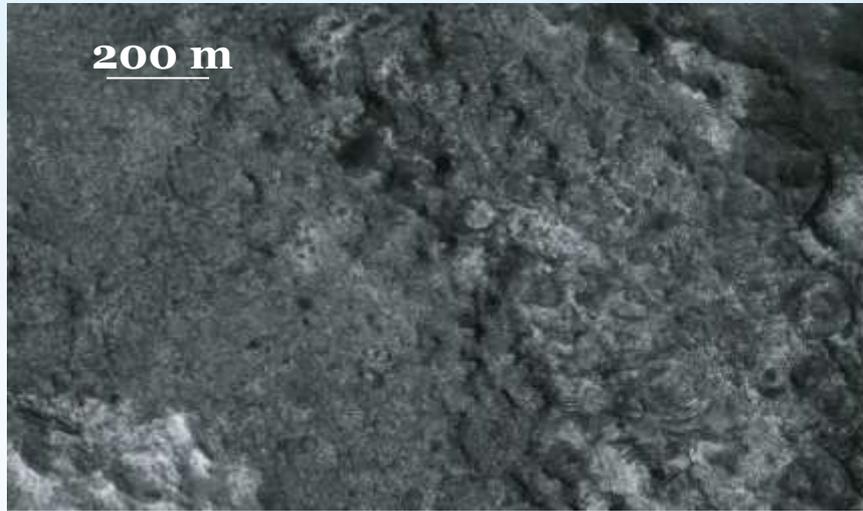


ROI: Pitted layers





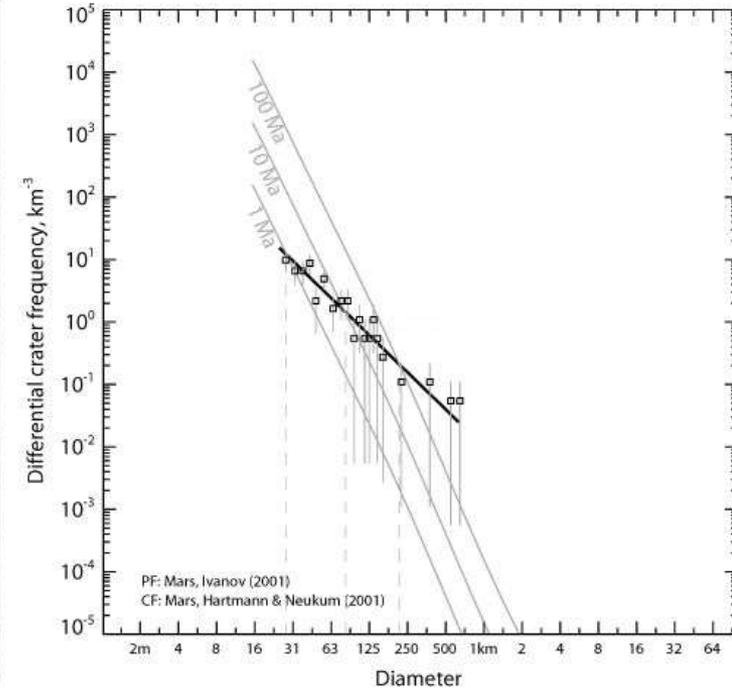
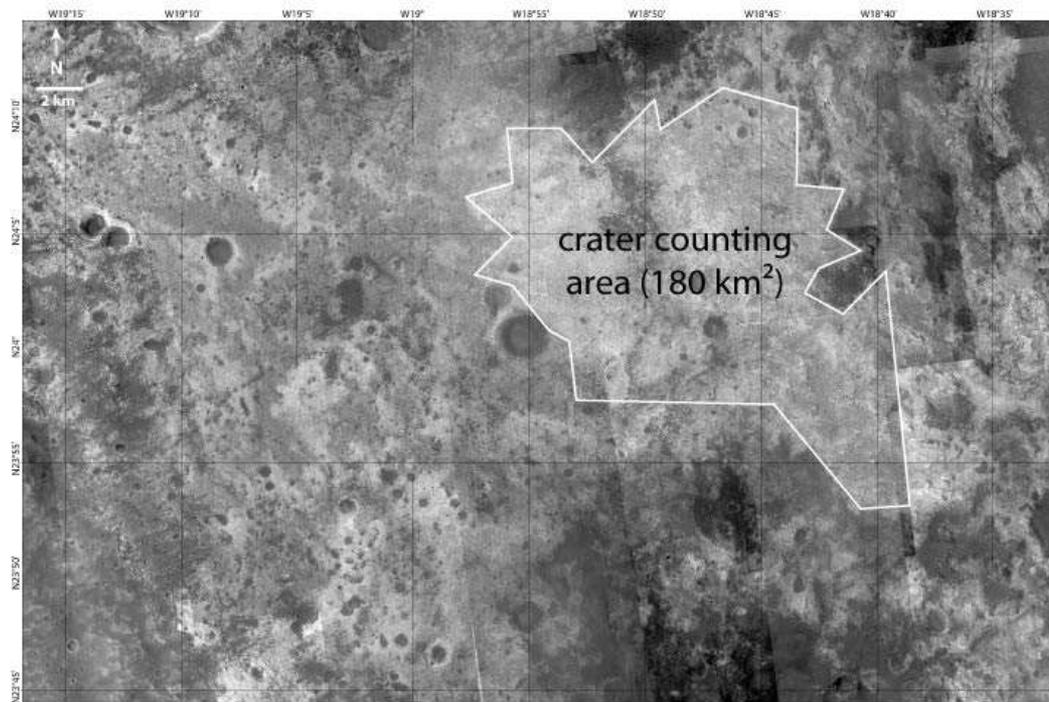
ROI: Paleo-features



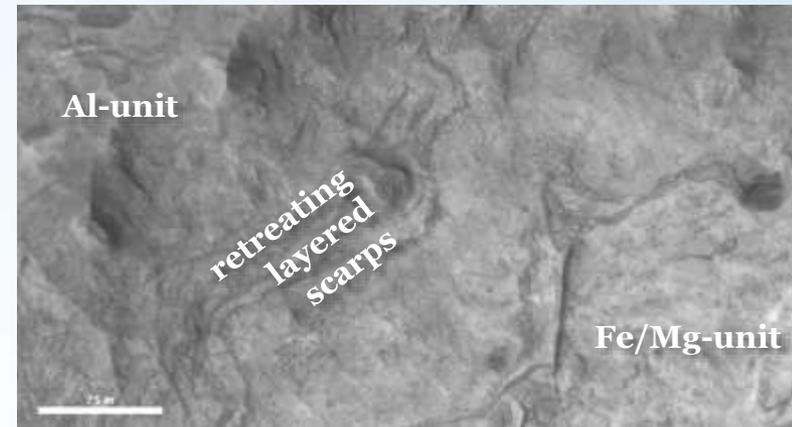
TARGETS & ROIs

Unit	Morphology	Mineralogy	Objectives	Ranking
Dunes	in local lows, generally ancient craters and valleys, direction: W-E in eastern part, SW-NE in western part	no hydrated signature, maybe basaltic sand if erosion product from the dark cap	...	5
Dark cap	flat-top mesas or crater filling, surrounded by dark talus: probably a "duricrust" protecting more easily erodible material	no hydrated signature, weak pyroxene signature (Loizeau et al., 2007)	Sample caching	4
Upper blue member	gentle slope forming unit, generally regular polygons 0.5-1.5 m across*	Al-phyllsilicates (montmorillonite, kaolinite), hydrated silica, ferrous component at the contact with the red unit	Habitability, biosignatures, sample caching	1
Red member (above paleo-sand-sheet)	generally irregular polygons 2-5 m across, with variation in fracture styles, few exposed layers*	Fe-smectite, ferrous component at the contact with the upper blue unit	Habitability, biosignatures, sample caching	1
Bright circular features	similar to red unit, but with many quasi-circular features 15-50m across	Fe-smectite	Habitability, biosignatures, sample caching	3
Paleo-sand sheet	rough, linear features trending NW-SE (different direction than surface dunes)*	no hydrated signature	Sample caching	3
Red member (on the plateau, below the paleo-dunes)	generally irregular polygons 2-5 m across, with variation in fracture styles, few exposed layers*, very eroded areas	Fe-smectite	Habitability, biosignatures, sample caching	2
Red member (on Mawrth Vallis flank and floor)	generally irregular polygons 2-5 m across, with variation in fracture styles*, with many exposed layers	Fe-smectite	Habitability, biosignatures, sample caching	2
Lower blue member	partly fractured in blocks up to 1.5 m across, exposed layers	possibly bassanite, or zeolite (spectrally close minerals when weak signal)	Habitability, biosignatures, sample caching	2
Dark paleo-surface	hummocky terrain with many quasi-circular features 10 to 100 m across, interpreted as exhumed filled craters	Fe-smectites	Habitability, biosignatures, sample caching	2
Putative Ancient Valleys	Part of ancient valleys, or inverted valleys, or a series of buttes that could be an ancient inverted valley	Eroded inside the clays, often filled by dark cap	Habitability, biosignatures, sample caching	2

PRESERVATION: RECENT EXPOSURE



- **very few small craters**
- **widespread aeolian activity and erosion due to dark sands**
- **easily erodible, very fine grained deposits**



Ellipse Plateau :To Summarize

- **Very diverse aqueous alteration** (from low to strong) inside the ellipse
- **Long and continuous stratigraphic altered** section
- **Diverse lithologies** that capture multiple environments (deposition, alteration & erosion)
- Reducing conditions, silica and very high clay content - **high preservation potential**
- **Fresh surfaces** (continuous erosion of clays) and **no diagenetic** overprinting
- Sample biosignatures in the **habitable environment** in which they were formed

Ellipse Plateau :To Summarize

- Consistent with a **paleosol sequence** ending in a **wetlands-like environment**
- **Reducing terrains** (soils) cause immediate preservation and can lead to **concentrated organics** if a biological cycle is active
- Unique window into **climate of early Mars (EN to H)**
- Dateable ~EH surface (**mafic capping unit**) present at numerous locations inside the landing ellipse

Ellipse Plateau :To Summarize

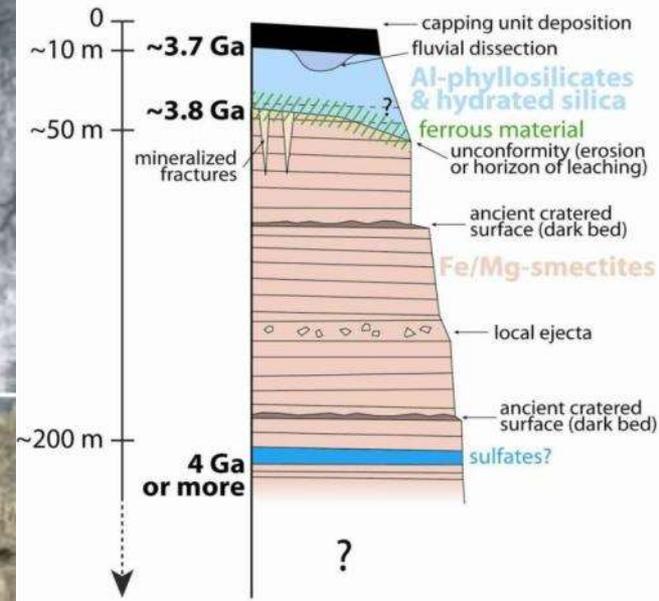
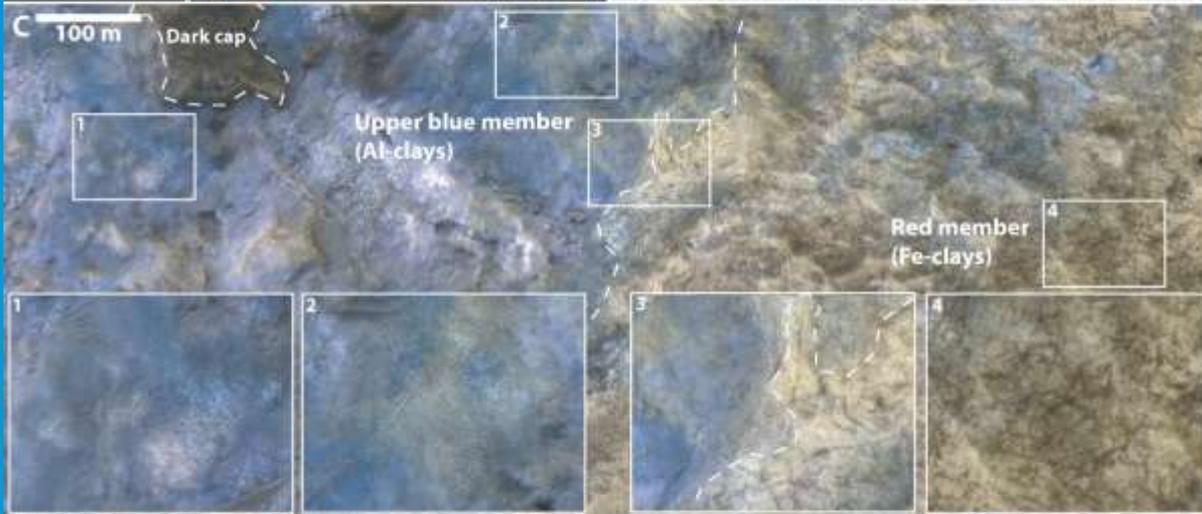
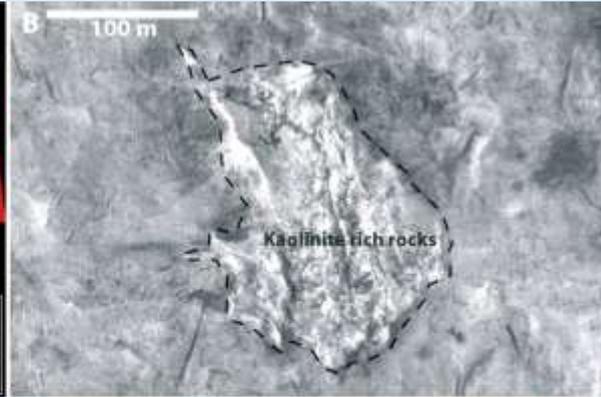
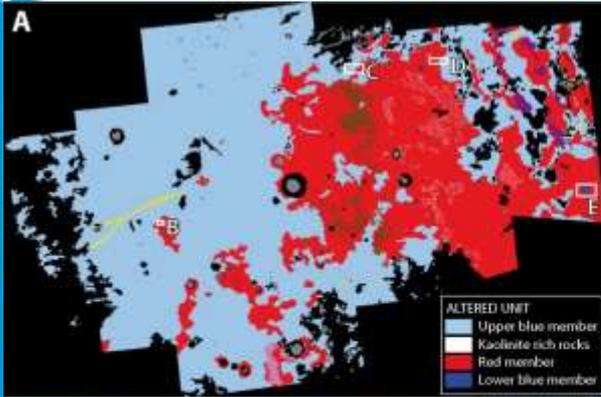
- Rubric is supported by tens of papers

Landing Site Factor	Mars 2020 Mission and Decadal Priority Science Factors																						
	Environmental Setting for Biosignature Preservation and Taphonomy of Organics							Type 1A & 1B Samples: Aqueous Geochemical Environments Indicated by Mineral Assemblages							Type 2 Samples: Igneous		Context: Martian History Sampled, Timing Constraints						
	Deltaic or Lacustrine (perennial)	Lacustrine (evaporitic)	Hydrothermal (<100°C) surface	Hydrothermal (<100°C) subsurface	Pedogenic	Fluvial/Alluvial	No diagenetic overprinting	Recent exposure	Crustal phyllosilicates	Sedimentary clays	Al clays in stratigraphy	Carbonate units	Chloride sediments	Sulfate sediments	Acid sulfate units	Silica deposits	Ferric Ox./Ferrous clays	Igneous unit (e.g. lava flow, pyroclastic, intrusive)	2nd Igneous unit	Pre- or Early-Noachian Megabreccia	Oldest stratigraphic constraint	Youngest stratigraphic constraint	Stratigraphy of units well-defined
Mawrth Vallis plateau	○		?	●	●	●	●	?	●	●			○	●	●	●	●	○		EN	EH	●	●

- ⇒ Perfect site to address in detail the questions of habitability & the potential origin and evolution of life on Mars
- ⇒ Minimizes the quantity of driving to maximize the quantity of field characterization and coring

ALTERED UNIT, *priority 1*

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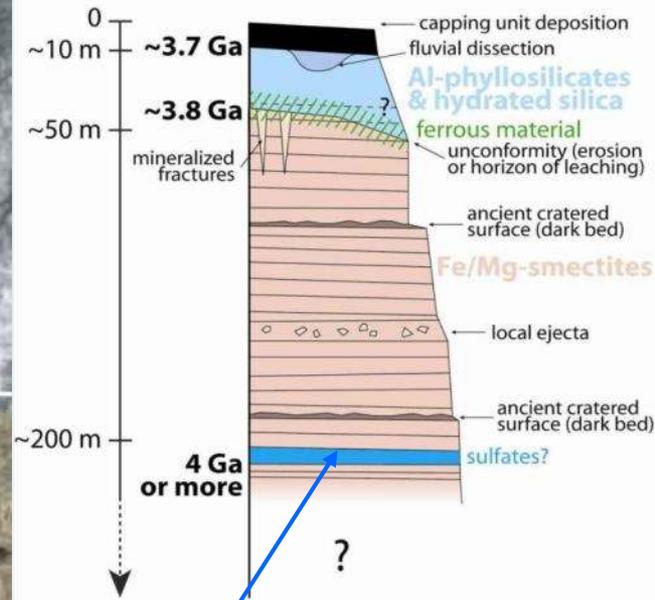
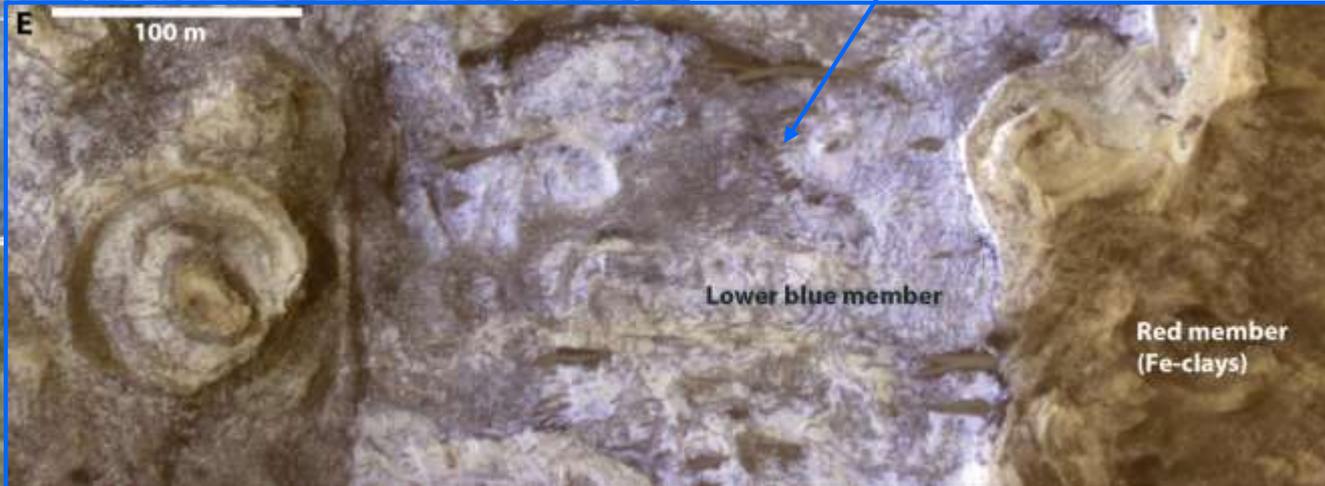
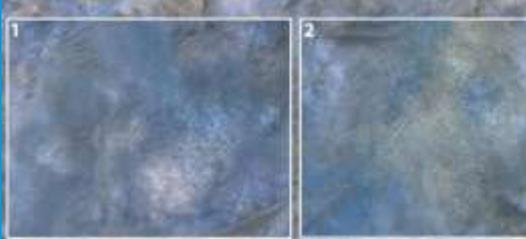
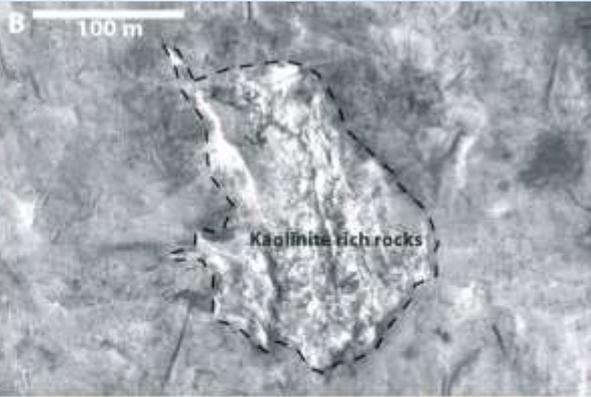
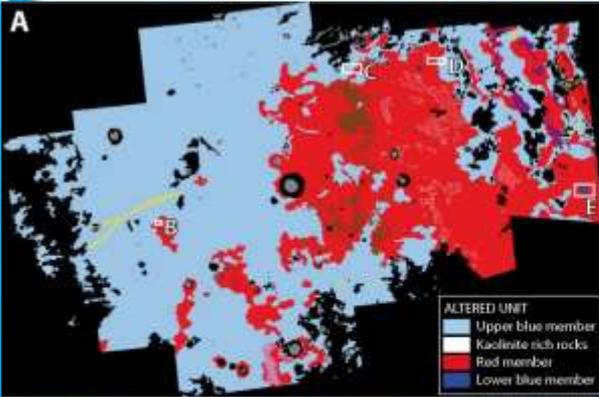
Transition from the Upper-blue to Red member

Highest priority

The alteration is also revealed by **difference in color and texture**

ALTERED UNIT, priority 1

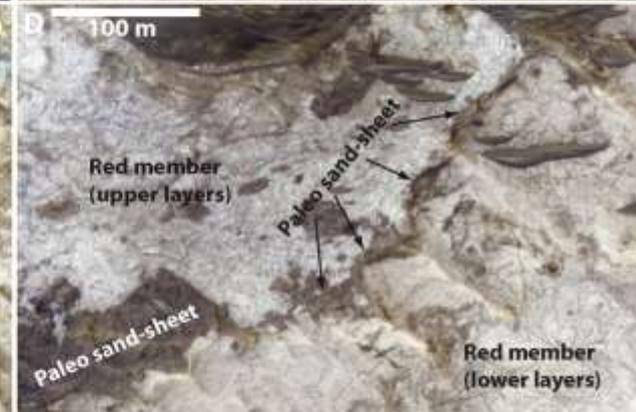
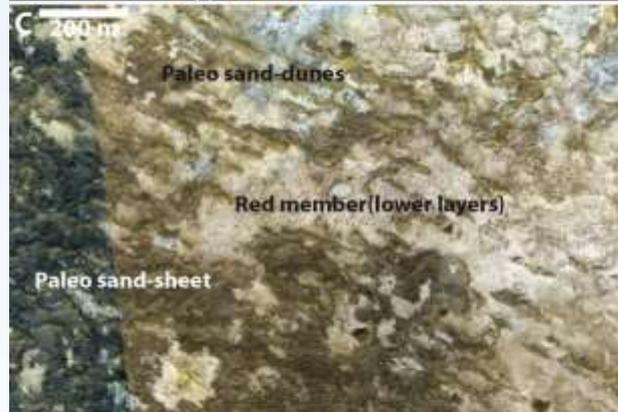
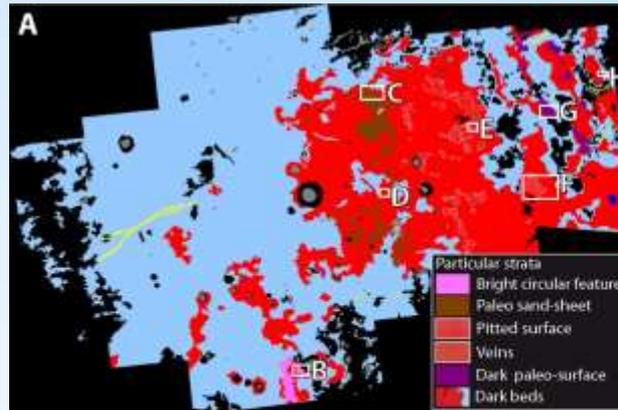
MARS 2020 at MAWRTH VALLIS



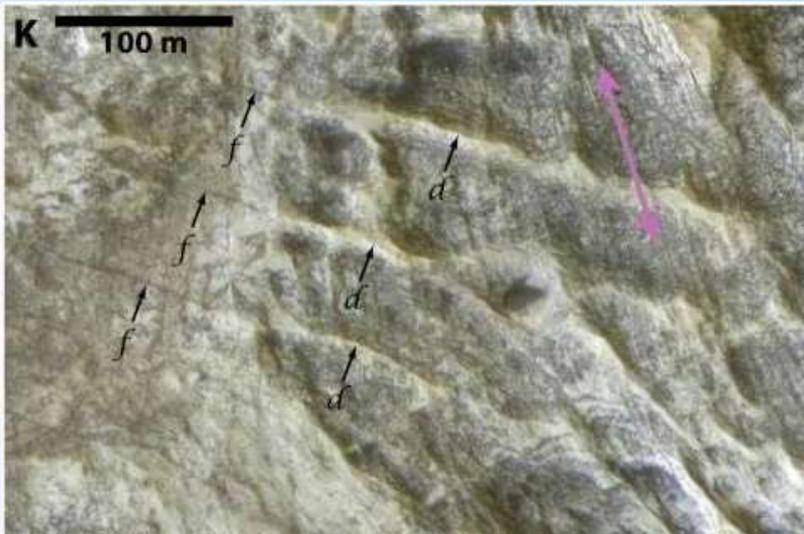
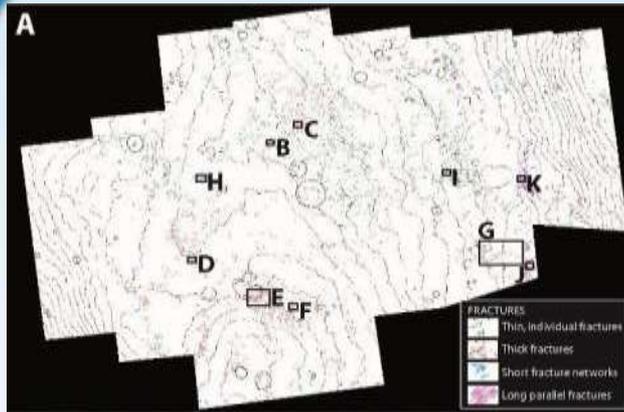
SPECIAL STRATA, *priority 2*

Special strata in the Red member

- rounded deposits
- paleo-sand sheet (parallel dunes)
- pits differential erosion
- veins differential erosion

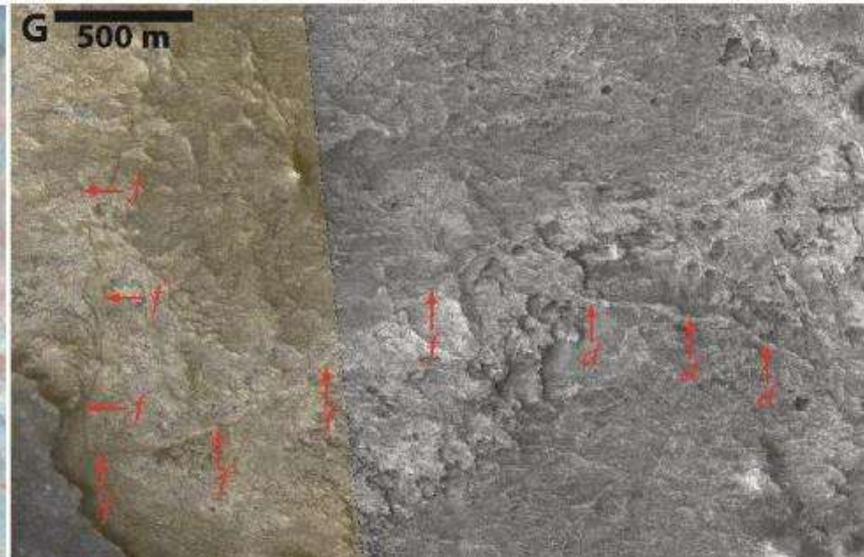
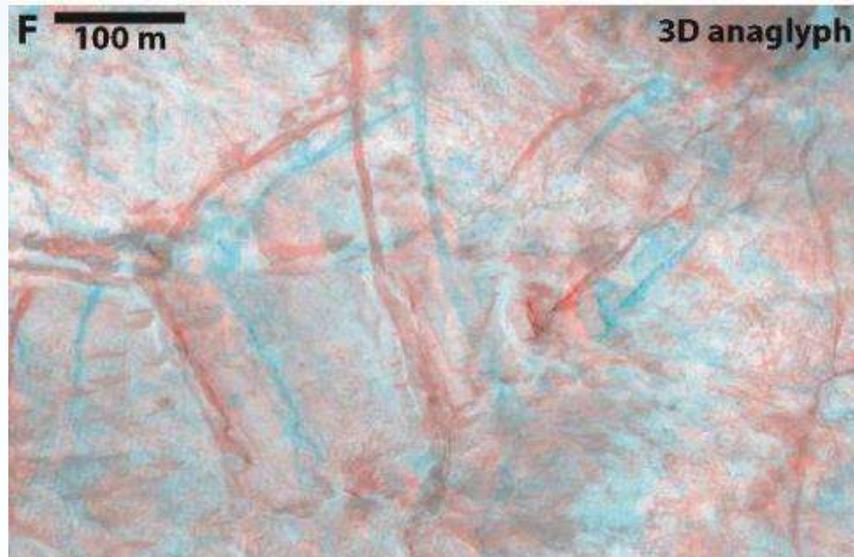


FLUID CIRCULATION, priority 2



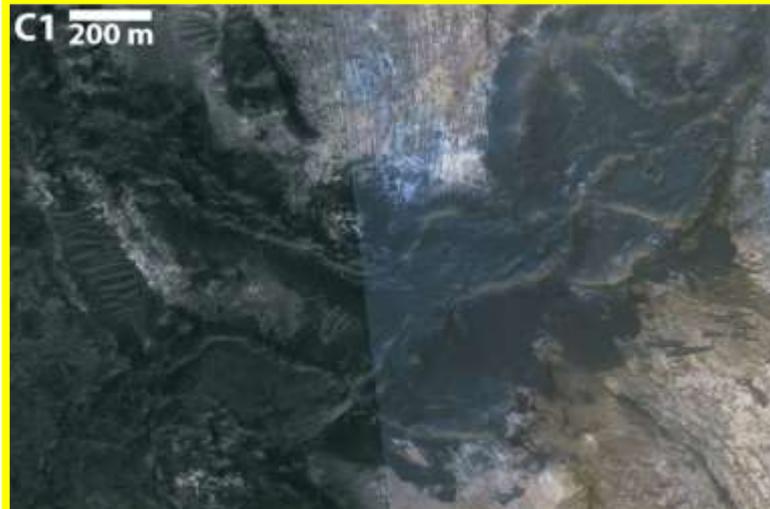
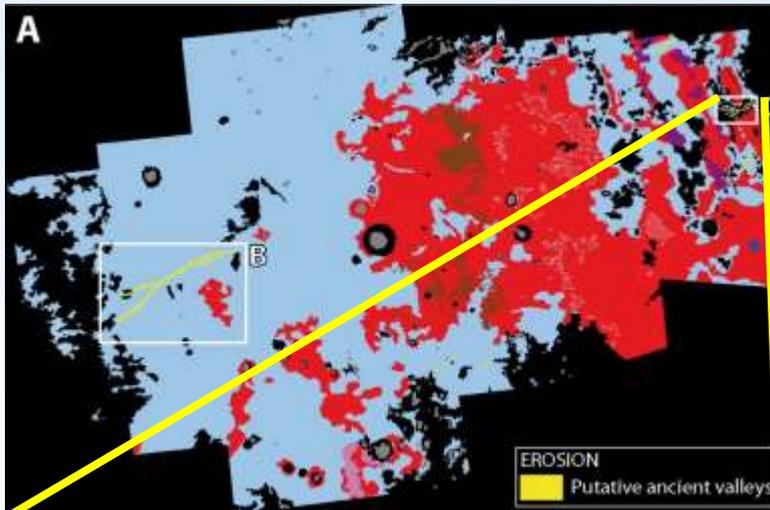
- halo-bounded fractures
- veins
- dense networks of fractures

=> alteration in the subsurface through fluid circulation



INVERTED VALLEYS, *priority 2*

- Fluvial features are present throughout units but are only obvious below capping unit due to inverted relief in resistant unit
- Relief inversion shows the softness of the clays wrt the dark cap



CAPPING UNIT, *priority 3*

- dateable **igneous** surface, regionally extensive
- eroded and redeposited (tallus, layered filling, dunes)
- rests of ejecta (15 km fresh crater south of the site)

MARS 2020 at MAWRTH VALLIS

